## 千葉大での研究内容

- 1. 90Sr Counter 開発研究
- 2. J-PARC E36 実験

2017. 10. 03 伊藤博士

#### <u>背景</u>

- 福島原発後、福島漁業が未だ再開しない、原因の一つは90Sr
- 90Srは骨に蓄積するため137Csと比較して摂取は危険
- 新鮮な海産物の短時間での<sup>90</sup>Sr放射能濃度測定が要求

#### <u>目的</u>

- 1時間で数十Bq/kgの90Srを測定するために低バックグラウンド
- 40Kからのβ線や宇宙線μ由来の雑音を抑制
- 大面積の有効面積

# Chapter 5

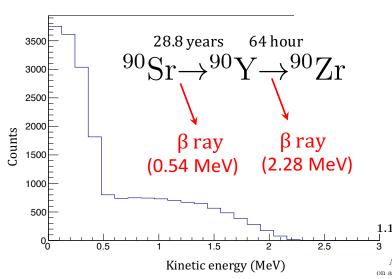
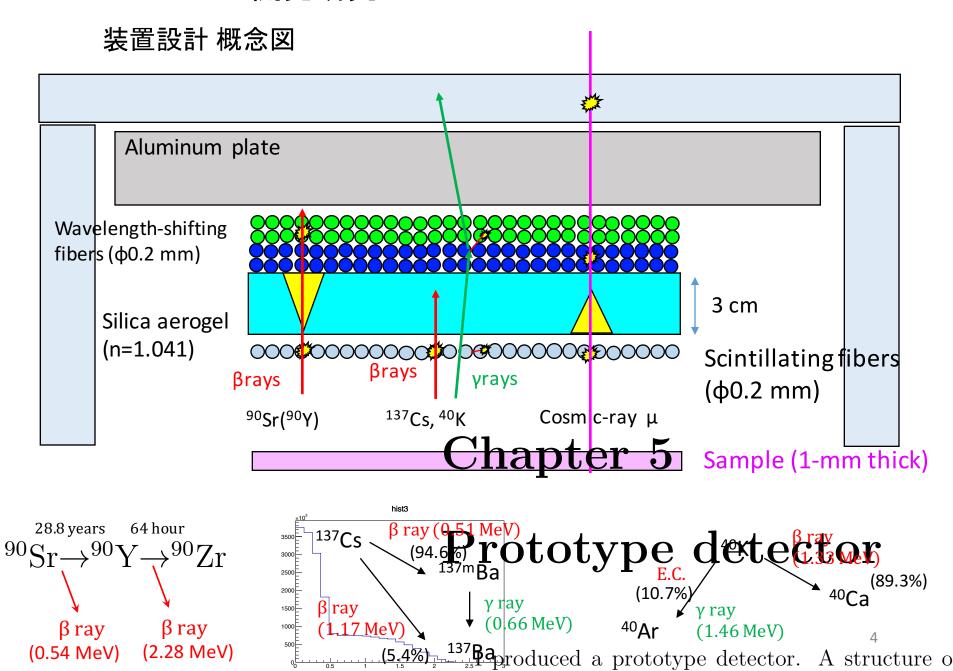


Fig. 1.1: Energy spectrum of  $\beta$  ray from  $^{90}\mathrm{Sr}$  and  $^{90}\mathrm{Y}$ . 2017/10203

Dose coefficient by ICRP publ.

A conventional method of measurement of "Sr Concentration is based on a chemical extraction after a sample was burned and became to ash and it takes usually few recks of an amount of the same of the

using a calorimeter or a magnetic spect: the end point in the presence of cosmic



## これまでの研究業績

### 1. 90Sr Counter 開発研究

#### 閾値型チェレンコフ検出器

<sup>40</sup>K からの β 線でチェレンコフ放射しない屈 折率の閾値は

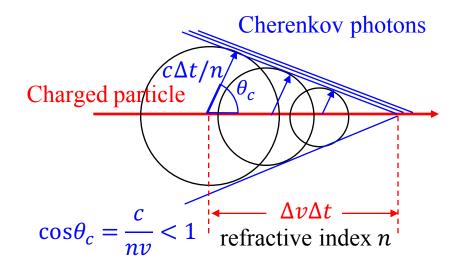
$$n_{\rm th} = \frac{m_{\rm e}c^2 + K}{\sqrt{(m_{\rm e}c^2 + K)^2 - m_{\rm e}^2c^4}}$$

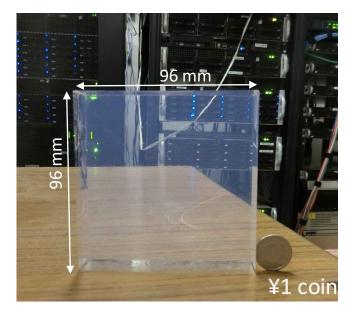
#### me は電子質量

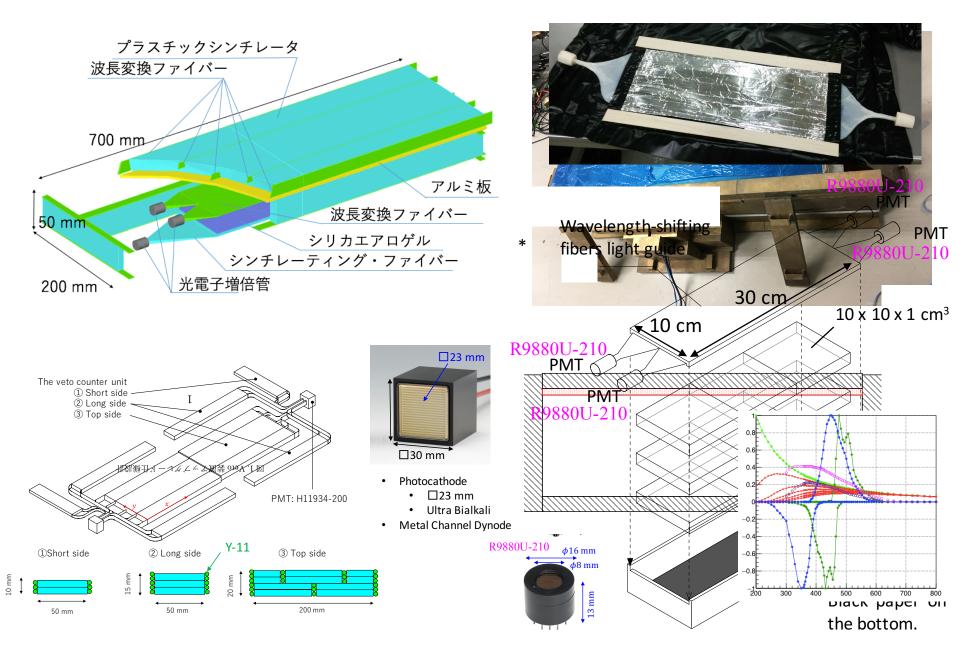
 $K=1.31~{\rm MeV}$ なので $n_{\rm th}=1.041$ この屈折率では $E\gamma=1.53~{\rm MeV}$ 未満の $\gamma$ 線がコンプトン散乱して反跳した電子もチェレンコフ放射しない。

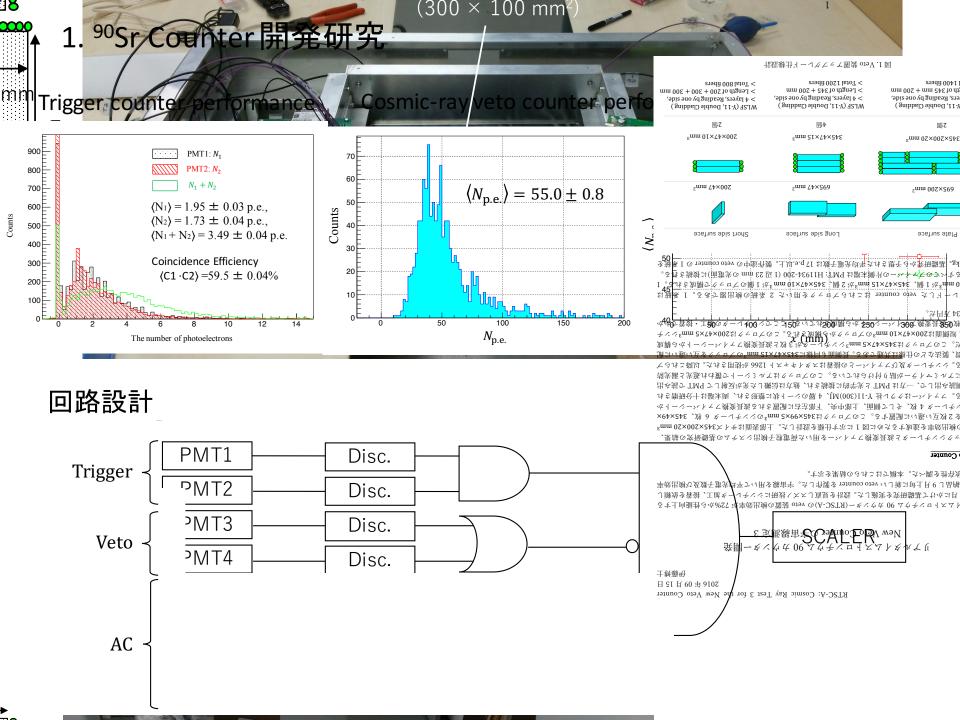
#### シリカエアロゲル

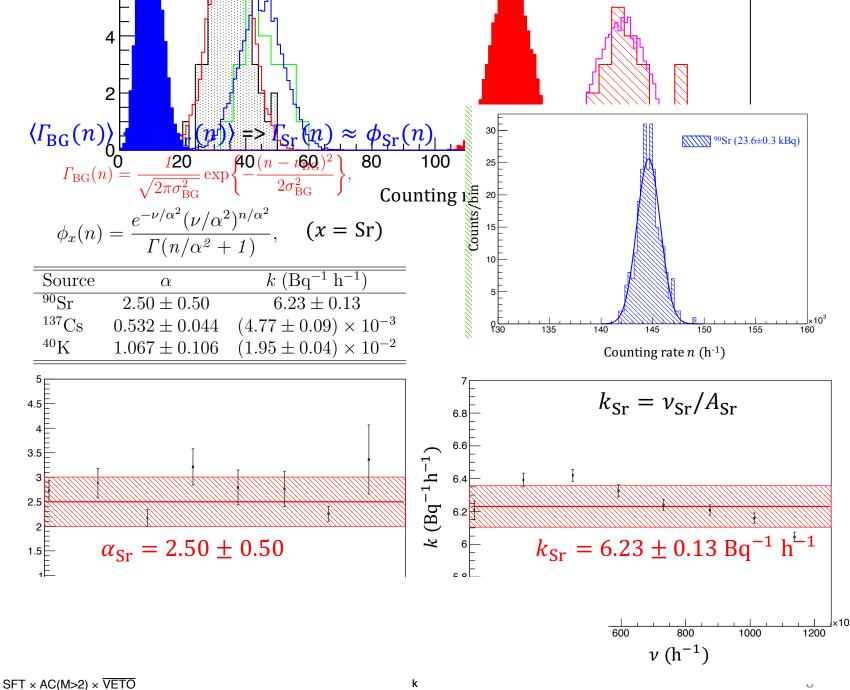
- SiO2と空気の混合体のように振る舞う個体
- 1.041のような屈折率を実現







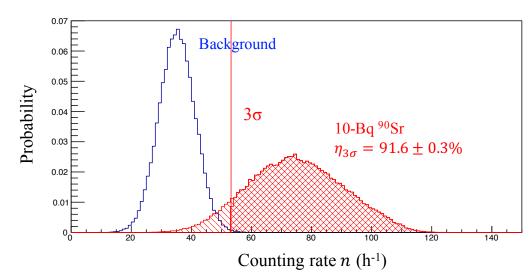


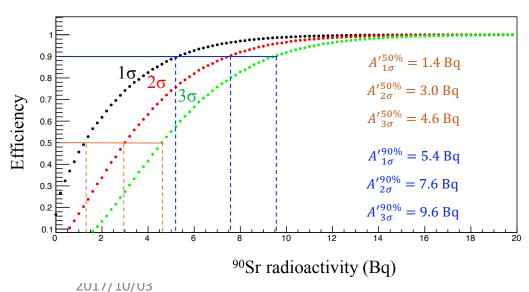


 $\boldsymbol{z}$ 

 $F \longrightarrow Y = 10 \text{ mm} \longrightarrow Y = 50 \text{ mm} \longrightarrow Y = 90 \text{ mm}$ 

#### 線源を用いた性能評価測定





- These curves show relations between  $^{90}$ Sr radioactivity and the efficiency for 1, 2, 3 $\sigma$  threshold condition.
- Typical detection limit is determined to be  $A_{3\sigma}^{50\%}$  satisfying  $\langle \Gamma_{\rm Sr}(n) \rangle > \langle \Gamma_{\rm BG}(n) \rangle + 3\sigma$ .
- $A_{3\sigma}^{50\%} = 4.6$  Bq at 1-hour measuring.

乾燥した海産物サンプルの密度1 g/cm<sup>3</sup>、体積圧縮率  $\epsilon$ =0.3、厚さ1 mm、質量 m=30 gとすると1時間測定における検出限界は、

$$A_{3\sigma}^{50\%} m \, \varepsilon^{-1} = 46 \, \text{Bq/kg}$$

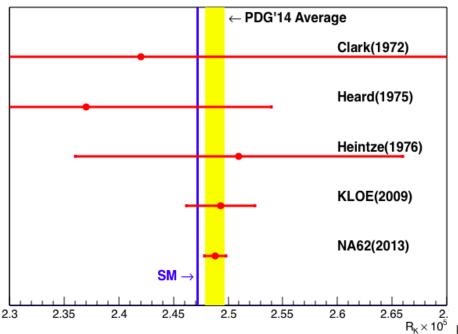
 $A \propto S^{-1}$ ; it expected to be 8.4 Bq/kg @S=1 m<sup>2</sup>

海水における検出限界は

$$A_{3\sigma}^{50\%} m \, \varepsilon^{-1} = 1.5 \, \text{Bq/L}$$

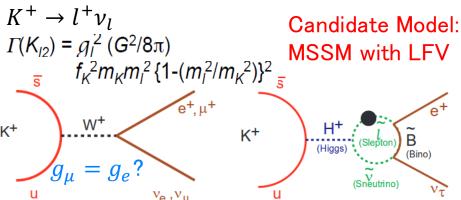
静止 $K^+$ を用いた $\Gamma(K^+ \to e^+ \nu_e)/\Gamma(K^+ \to \mu^+ \nu_\mu)$ の精密測定実験

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	$R_K \times 10^5$	$\Delta R_K/R_K$
KLOE(2009)	$2.493 \pm 0.025 \pm 0.019$ (stat) (sys)	1.26%
NA62(2013)	$2.488 \pm 0.007 \pm 0.007$ (stat) (sys)	0.40%
SM	$2.477 \pm 0.001$	0.04%
Initial goal of		0.25%

$$\begin{split} R_K^{SM} &= \frac{\Gamma(K^+ \to e^+ \nu_e)}{\Gamma(K^+ \to \mu^+ \nu_\mu)} \\ &= \frac{m_e^2}{m_\mu^2} \Big( \frac{m_K^2 - m_e^2}{m_K^2 - m_\mu^2} \Big)^2 \frac{(1 + \delta_r)}{\text{radiative correction}} \\ &\quad \text{helicity suppression} \end{split}$$

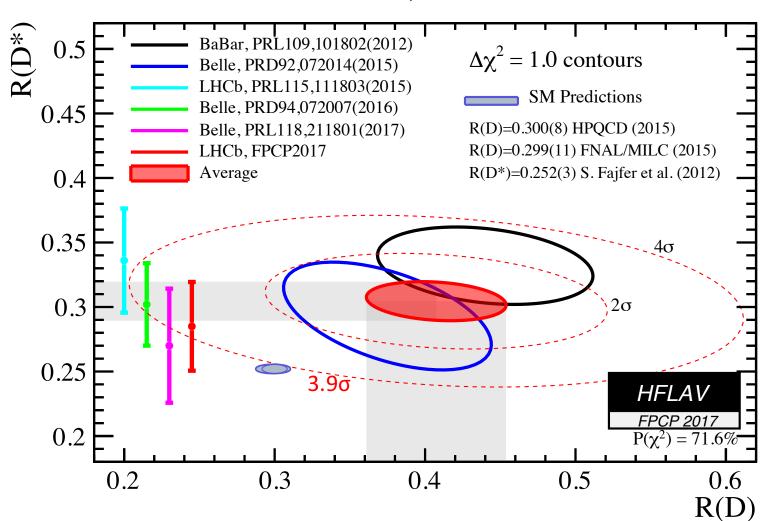


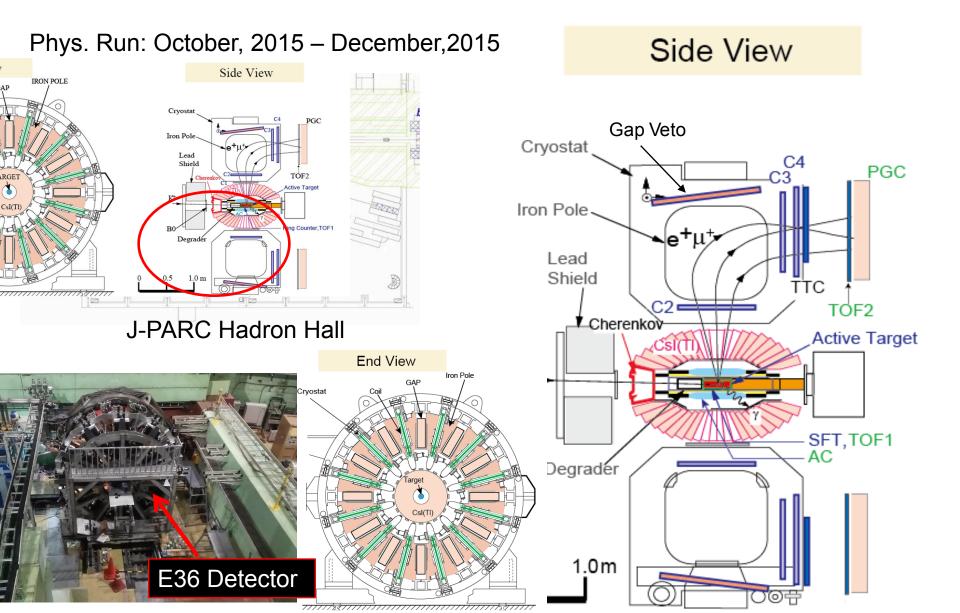
$$R_K^{LFV} = R_K^{SM} \left( 1 + \frac{m_K^4}{M_{H^+}^4} \cdot \frac{m_\tau^2}{m_e^2} \Delta_{13}^2 \tan^6 \beta \right)$$
$$\sim R_K^{SM} (1 + 0.013_{\text{max}})$$

Phys. Rev. D **74** 

$$\mathcal{R}(D) = \frac{\mathcal{B}(\overline{B} \to D\tau^{-}\overline{\nu}_{\tau})}{\mathcal{B}(\overline{B} \to D\ell^{-}\overline{\nu}_{\ell})}, \quad \mathcal{R}(D^{*}) = \frac{\mathcal{B}(\overline{B} \to D^{*}\tau^{-}\overline{\nu}_{\tau})}{\mathcal{B}(\overline{B} \to D^{*}\ell^{-}\overline{\nu}_{\ell})}$$

where I refers to either an e or  $\mu$ .

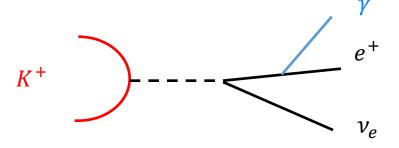


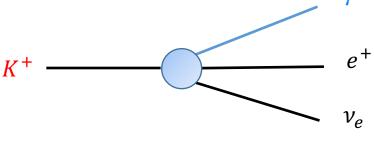


$$R_K^{SM} = \frac{\Gamma(K_{e2}) + \Gamma(K_{e2\gamma}(IB))}{\Gamma(K_{\mu 2}) + \Gamma(K_{\mu 2\gamma}(IB))}$$

Background:  $Ke2\gamma$  (SD)

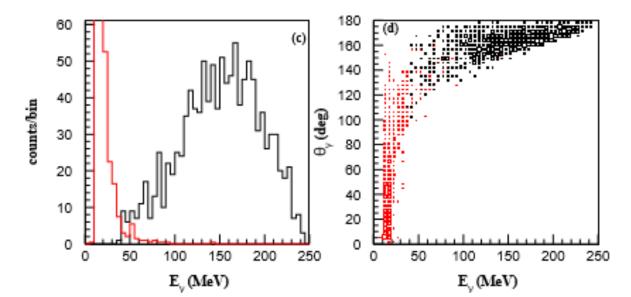
 $Ke2\gamma:K^+\to e^+\nu_e\gamma$ 





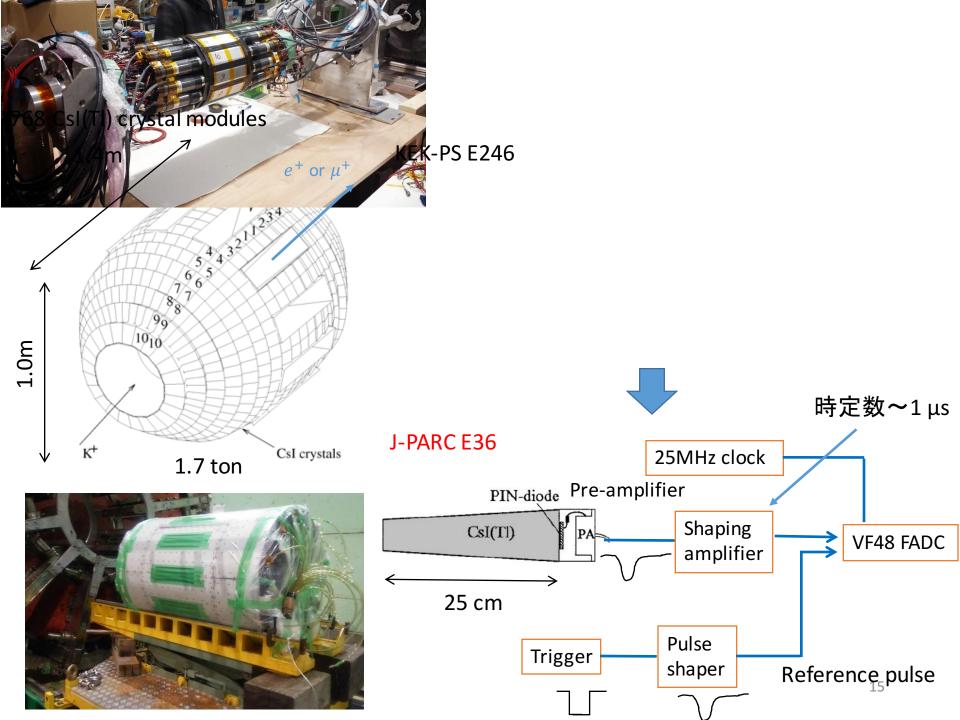
#### Internal Bremsstrahlung (IB)

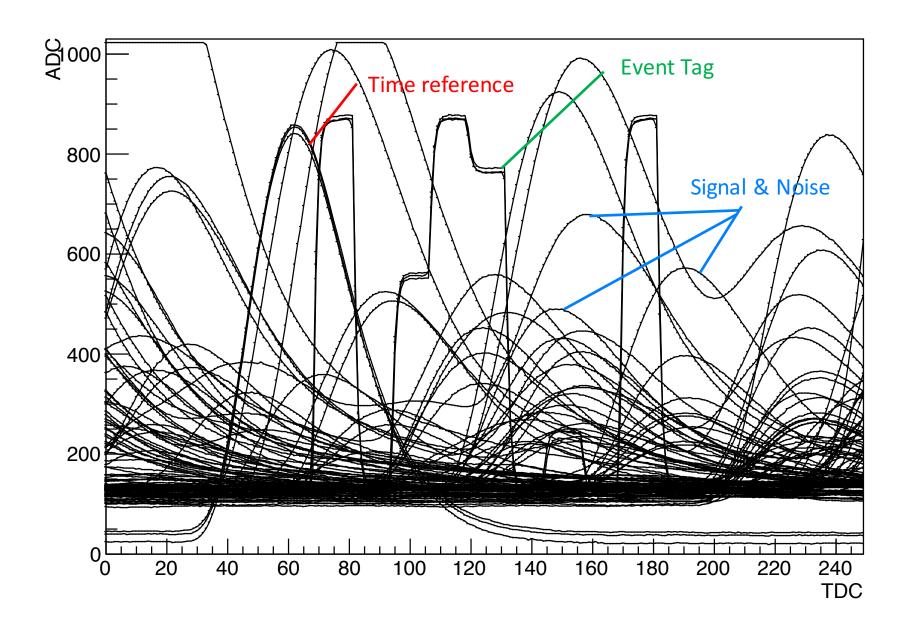
Structure Dependent (SD)



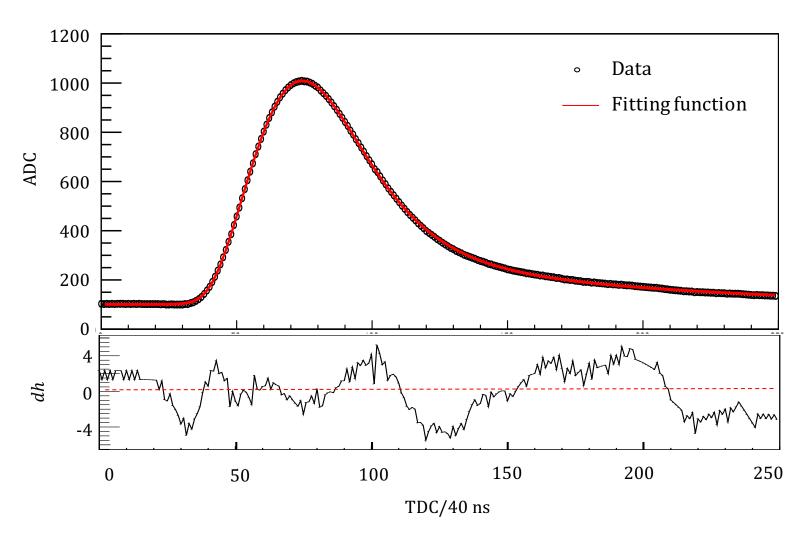
 $\Gamma(\text{Ke2}\gamma(\text{SD})) \sim 9.4 \times 10^{-6}$ 

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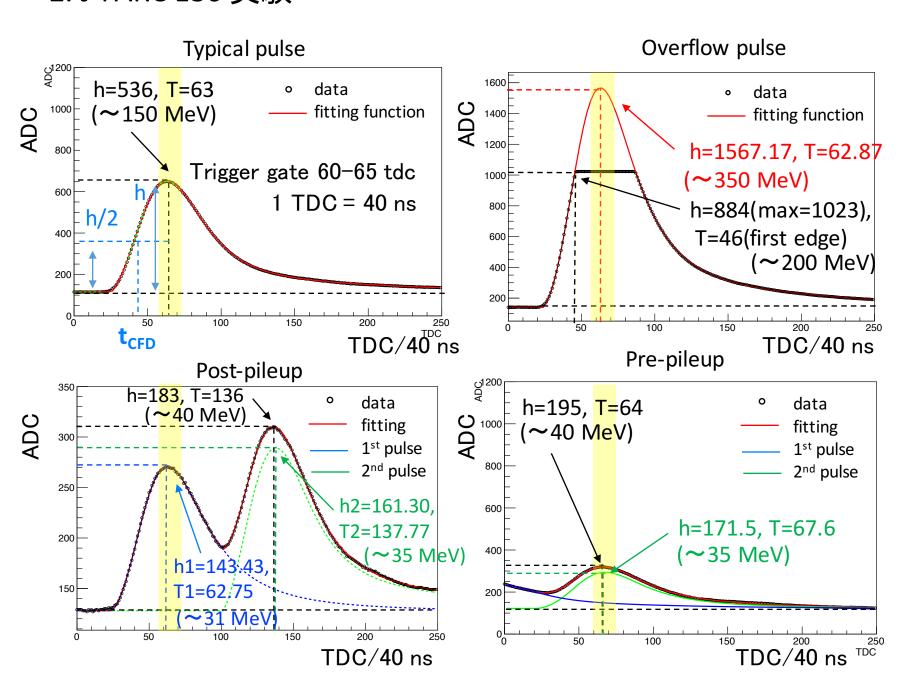




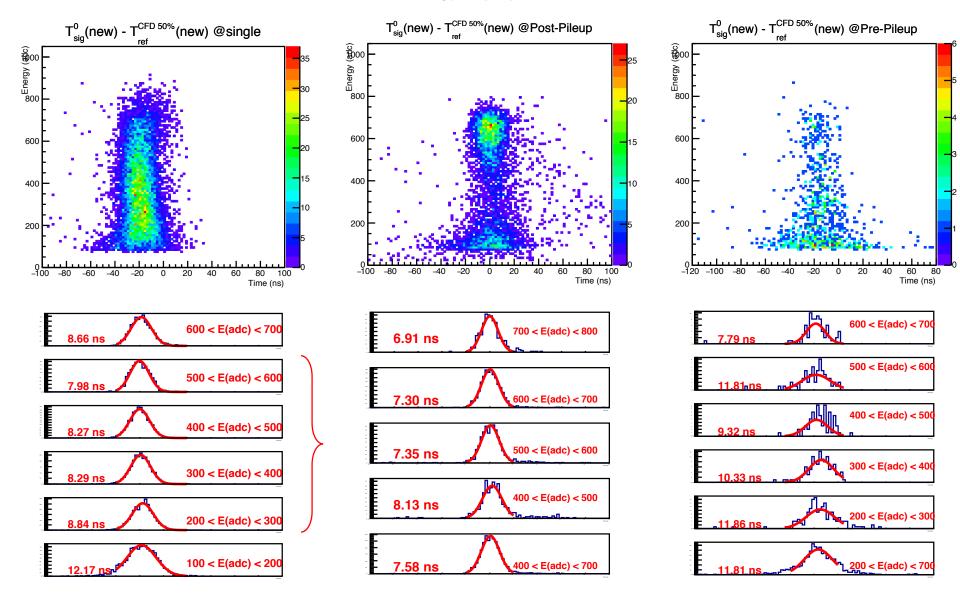
## Development of Waveform Model



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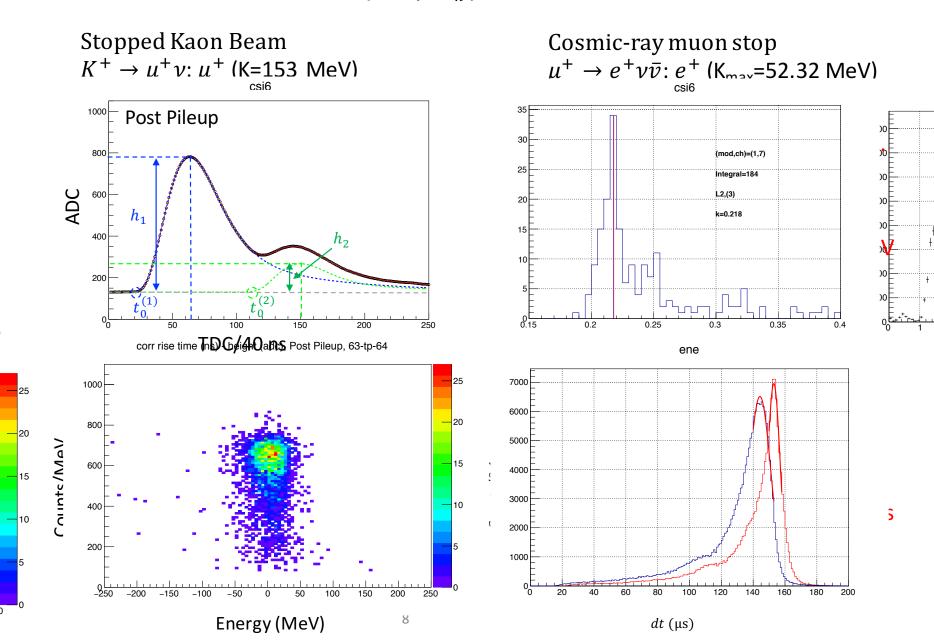


#### 時間分解能評価

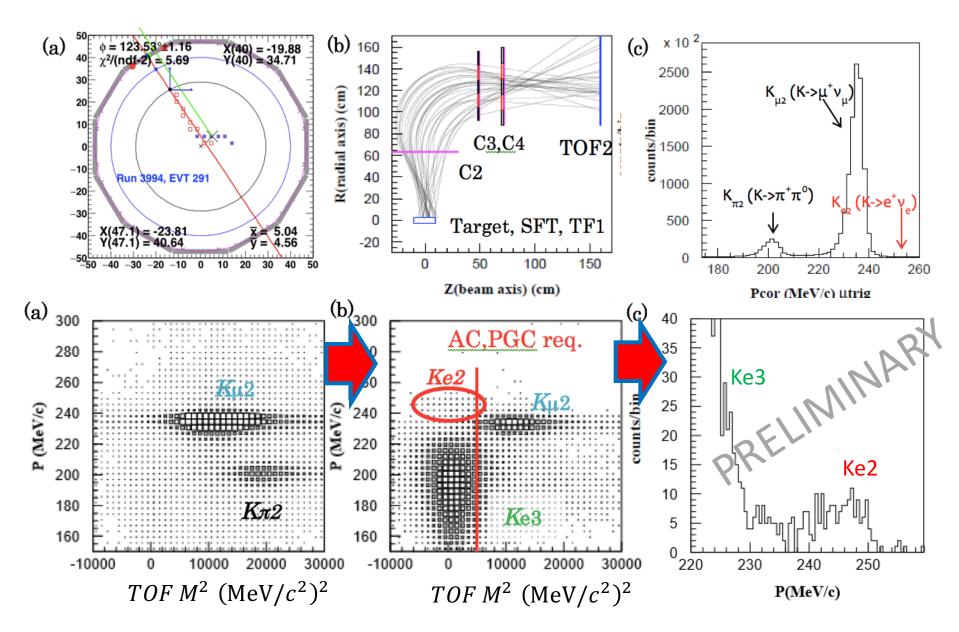


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#### エネルギー較正

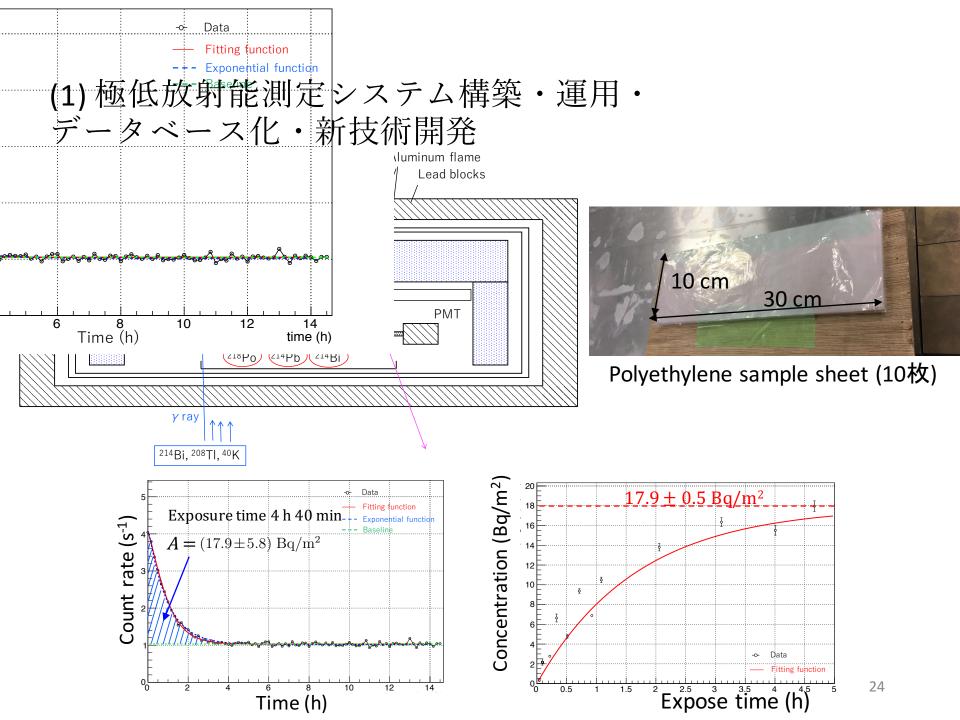


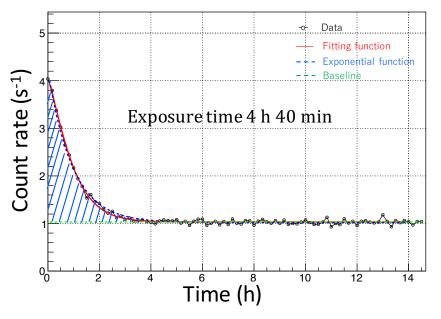
#### Ke2γ Background study

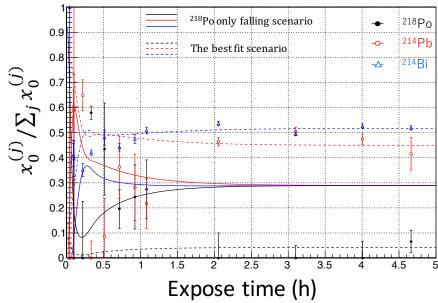


#### I DV DC E36 卓粋 dalitz angle2 angle ang Ee > 229 MeV 100 100 80**F** At birth 80 20**F** 150 50 Egamma Egamma Egamma Eγ (MeV) **Εγ (Men/g)**e3 enel ene Cogsts/bin 2500 1600 1400 2000 1200 1000 1500 800 1000 600 400 E 500 200 100 250 El (MeV) 120 140 160 180 Opening angle (deg.) 50 200 250 Egamma (MeV) 150 200 100 150 CsI Eγ (MeV) CsI Eγ (MeV) 22

# Buck up







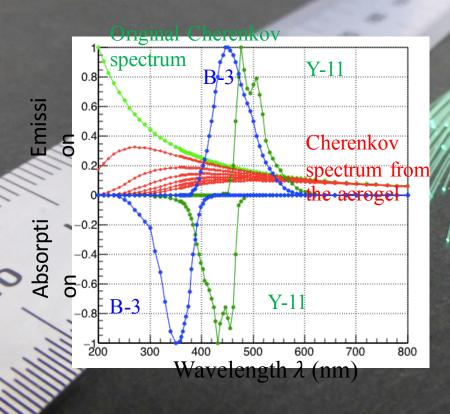
Isotope j = 1, 2, 3 are <sup>218</sup>Po, <sup>214</sup>Pb, <sup>214</sup>Bi, respectively.

 $\lambda_j = \tau_j^{-1}$ : inverse of life time of isotope j $x_0^{(j)}$ : initial intensity of isotope j

 $R_{BG}$ : background rate

# これまでの研究業績

## 1. 90Sr Counter 開発研究



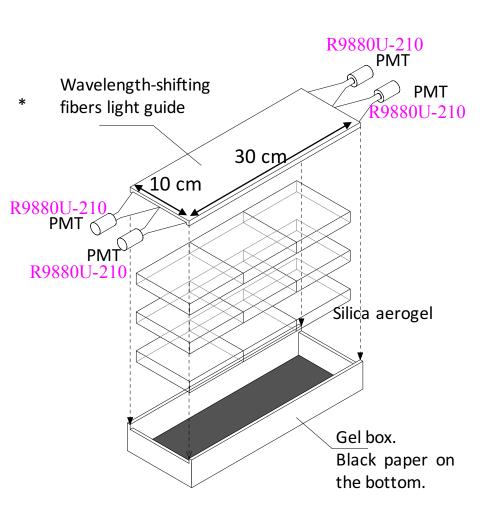
Double cladding φ0.2 mm

Kuraray Y-11(300) MJ

B-3(300)MJ

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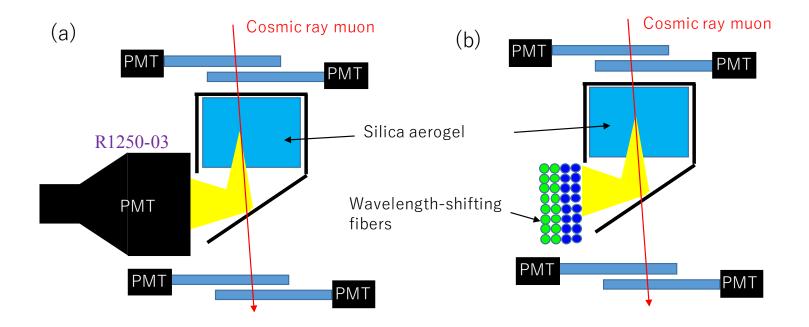






# 2. Basic concept an aerogel Cherenkov detector using wavelength-shifting fibers

Measurement of light collection efficiency



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 $P(k,\nu) = e^{-\nu}\nu^k/k!$ 

 $\langle N_{p.e.} \rangle (= \nu)$ 

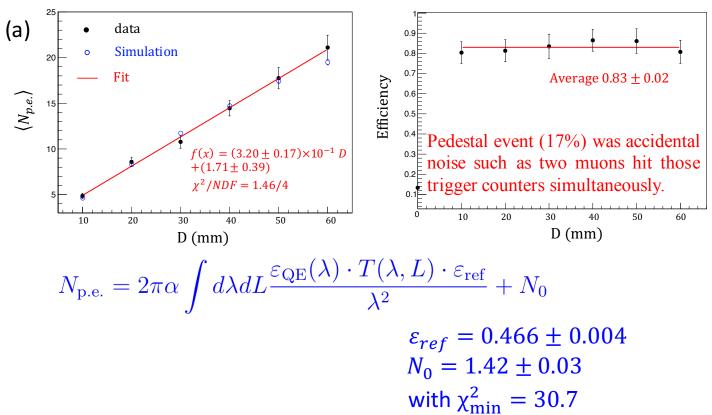
# 2. Basic concept an aerogel Cherenkov detector using wavelength-shifting fibers Poisson Function Fitting

Measurement of light collection efficiency

D = 10 mm $\langle N_{n.e.} \rangle = 4.9 \pm 0.2$ Cosmic ray muon (a)  $\chi^2/NDF = 15.5/19$ PMT histb2 histb3 R1250-03 22 20 18 16 14 12 10 8 6 6 4 4 D (mm) N<sub>p.e.</sub> w<sub>n.e.</sub> PMT D = 50 mmD = 60 mm $\langle N_{ne} \rangle = 17.8 \pm 0.4$  $\langle N_{ne} \rangle = 21.1 \pm 0.5$ /NDF = 51.7/49/NDF = 52.0/55201,, 10,00  $N_{\rm p.e.}$ 

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Measurement of light collection efficiency



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