Development of an alpha-particle imaging detector based on a low radioactive micro-time-projection chamber

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Abstract

An important issue for rare-event-search experiments, such as the search for dark matter or neutrinoless double beta decay, is to reduce radioactivity of the detector material and the experimental environment. The selection of materials with low radioactive impurity, such as isotopes in the uranium and thorium chains, requires a precise measurement of surface radioactivity. An alpha-particle detector has been developed based on a gaseous micro-time-projection chamber. A low-*α µ*-PIC with reduced alpha-emission background was installed in the detector. The detector offers the advantage of position sensitivity, which allows the alphaparticle contamination of the sample to be imaged and the background to be measured at the same time. The detector performance was measured by using an alpha-particle source. The measurement with a sample was also demonstrated and the sensitivity is discussed.

Keywords: Alpha-particle detector, Position sensitivity, Time projection chamber, μ -PIC, Low background

¹ **1. Introduction**

 Approximately 27% of the universe is domi- nated by nonbarionic matter, called dark mat- ter. Although many experimental groups have been searching for dark matter, no direct detection of dark matter has yet been reported. Typical experiments that search for dark matter are performed by using massive, low-background detectors. Although the DAMA group has observed the annual modula-10 tion with a significance of 9.3σ as the dark matter contribution [1], other groups such as XENON1T [2] and LUX [3] did not reproduced the signal. Meanwhile, a direction-sensitive method has been focused because of an expected clear anisotropic signal due to the motion of the solar system in the galaxy [4]. The NEWAGE group precedes a three-dimensionally sensitive dark matter search with a

 micro-time-projection chamber (micro-TPC) and the main background is surface alpha particles from U and 232 Th in the detector material or in the μ -PIC [5].

 Neutrinoless double beta (0*νββ*) decay is a lepton-number-violating process, which suggests the neutrino as a Majorana particle (it is its own antiparticle) and provides the absolute neutrino mass. The GERDA [6] and KamLAND-Zen [7] 27 groups recorded a lower-limit half-life over 10^{25} yr ²⁸ at 90%CL by using ⁷⁶Ge and ¹³⁶Xe, respectively, and the 0*νββ* decay has yet to be observed. Con- versely, a tracking system for two electrons pro- vides strong evidence of the 0*νββ* decay process. The NEMO3 group precedes the measurement with at $T_{1/2} > 2.5 \times 10^{23}$ yr (90%CL) for ⁸²Se [8], and $T_{1/2}$ > (1.1 – 3.2) \times 10²¹ yr (90%CL) for ¹⁵⁰Nd [9] 35 and a contamination of 208 Tl and 214 Bi in the detec- tor dominates the background. The SuperNEMO group has developed the BiPo-3 detector to mea-

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 sure the radioactive impurities with sensitivity less ³⁹ than 2 μ Bq/kg (90%CL) for ²⁰⁸Tl and 140 μ Bq/kg 40 (90%CL) for ²¹⁴Bi [10]. Therefore, the background of 0*νββ* decay is not only a contamination by the end point of continuous energy in an ordinary 2*νββ* decay process, but also the radiative impurities such 44 as 238 U and 232 Th in the detector.

 To estimate the radioactive impurities in the de-⁴⁶ tector material, the XMASS group measured ^{210}Pb $_{47}$ and $_{210}$ Po in the bulk of copper by using a commer- cial alpha-particle detector (Ultra-Lo 1800, XIA) [11]. The alpha detector has a good energy resolu- tion (as explained in Sec. 3.2) and a mechanism to reduce the background by waveform analysis, and thus a sensitivity is $\sim 10^{-4} \alpha/cm^2/hr$. However, it has no position sensitivity. A sample such as a micro pattern gas detector board has not an uni- form radioactive contamination. For example, the impurities might be contaminated to the electrodes in a pattern making process. Therefore, a position- sensitive alpha detector is required to select mate-rials for the rare-event-search experiments.

 This paper is organized as follows. The details 61 of the alpha-particle detector, setup, low- α micro ϵ ² pixel chamber (μ -PIC), gas circulation system, elec- tronics, and trigger data acquisition system are de- scribed in Sec*.* 2. The performance check that uses the alpha-particle source, a sample test, and back- ground estimation are described in Sec*.* 3. The remaining background of the detector and future prospects are discussed in Sec*.* 4. Finally, the study is concluded in Sec*.* 5.

2. Alpha-particle imaging detector based on gaseous micro-TPC

 A new alpha-particle detector was developed based on a gaseous micro-TPC upgraded from the NEWAGE-0.3a detector [12] which was used to search for dark matter from September, 2008 to January, 2013. The detector consisted of the micro- TPC using a low-*α µ*-PIC, a gas circulation system, and electronics, as shown in Fig.1. The TPC was enclosed in a stainless-vessel for the gas seal during the measurement.

81 2.1. Setup and configuration

⁸² Figure 2 shows a schematic view of the detector, ¹¹⁰ where the gas volume is $(35 \times 35) \times 31$ cm³. The de-tector was placed underground at the Kamioka fa-

Fig. 1: Photographic of detector.

By the means to possess the cosmic system is a more of the cosmic of the cosmic containment
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solution. For cosmic the c An oxygen-free copper plate with a surface pol- $\frac{1}{87}$ ished to a roughness of 0.4 μ m was used as the drift plate. The drift plate had an opening with a size of $(9.5 \times 9.5 \text{ cm}^2)$ as a sample window. A copper mesh was set on the drift plate to hold the sample at the window area, as shown in Fig. 3. The electrons ionized by the alpha particles drift toward the *µ*-PIC with a vertical upward-pointing $_{94}$ electric field E . CF₄ gas, which was also used in the NEWAGE-0.3a, was used as the chamber gas because of the low diffusion properties. The pres- sure was set at 0*.*2 bar as a result of the optimiza- tion between the expected track length and the detector stability. The track length was expected to be longer, which improved the tracking perfor- mance when the gas pressures was low, while the discharge rate of the *µ*-PIC increased. The electric 103 field in the drift volume, $E = 0.4 \text{ kV/cm/bar}$, was formed by supplying a negative voltage of 2*.*5 kV and placing field-shaping patterns with chain resis- tors every centimeter [13]. The drift velocity was $107 \text{ } 7.4 \pm 0.1 \text{ cm}/\mu\text{s}$. The μ -PIC anode was connected to $+550$ V. The typical gas gain of μ -PIC was 10^3 at *∼* 500 V.

2.2. Low-*α µ*-PIC

The background study for the direction-sensitive dark matter search suggests that *µ*-PIC has radia-113 tivoactive impurities of U and 232 Th which emit

Fig. 2: Schematic cross section of detector setup.

 9.5×9.5 cm²) and copper support mesh. Fig. 3: Drift plate with a sample window (hole size is

 alpha particles [5]. A survey with a HPGe detec-¹¹⁵ tor revealed that μ -PIC's glass cloth was the main background source, and so the impurities were re- moved [14]. Details of the device with the new ma-118 terial, a low- α μ -PIC, will be described in Ref [15].

2.3. Gas circulation system

 A gas circulation system that uses activated char- coal pellets was developed for radon-background suppression and to protect a against gain deteri- oration due to the outgassing. A pump (EMP, MX-808ST-S) and a needle-type circulate meter (KOFLOC, PK-1250) were used to flow the gas at a 172 126 rate of $\sim 500 \text{ cm}^3/\text{min}$. The gas pressure was mon- itored to ensure the stable operation of the circula- tion system and as maintained within an increase of *∼* 2% for several weeks.

2.4. Electronics and trigger data acquisition system

 The electronics for the *µ*-PIC readout consisted of amplifier-shaper discriminators [16] for 768 anode and 768 cathode signals and a position-encoding module [17] to reconstruct the hit pattern. A data acquisition system consisted of a memory board to record tracks and a flash analog-to-digital con- verter (ADC) for the energy measurement. The flash ADC with 100 MHz sampling recorded the sum signal of the cathode strips with a full time range of 12 *µ*s. The anode sum signal issued the trigger. With this way of triggering, in contrast to the trigger by signal (for example, primary scintil- lation) in the TPC before the drift, the absolute position along the drift direction cannot be mea- sured. However, because the alpha particles were expected to be emitted from the sample, the drift- along coordinate of the emission point was assumed to be the position of the drift plate.

3. Performance check

3.1. Alpha-particle source

 $A = 10 \times 10 \text{ cm}^2$ copper plate with ²¹⁰Pb accu- mulated on the surface was used as an alpha- particle source for the energy calibration and energy-resolution measurement [11]. The source emits alpha particles with an energy of 5*.*3 MeV as a decay of 2^{10} Po. The alpha-particle emission rate (hereinafter called the *α* rate) of the source ¹⁵⁸ plate was calibrated to be $1.49 \pm 0.01 \alpha \text{ s}^{-1}$ for 4.8– 5*.*8 MeV by using the Ultra-Lo 1800 [11].

3.2. Energy calibration

 An energy calibration was conducted with the alpha-particle source. The energy was calculated from the flash ADC waveform. Figure 4 shows a typical energy spectrum of the alpha-particle source. The energy resolution was estimated to be $166 \quad 6.7\% \text{ (1}\sigma) \text{ for } 5.3 \text{ MeV}$, which is worse than the Ultra-Lo 1800 resolution of 4.7% (1*σ*) for 5.3 MeV. This deterioration was thought to be due to the gain variation of the *µ*-PIC detection area.

3.3. Event reconstruction

Figure 5 shows a typical event display with the tracks and flash ADC waveform data for alpha-173 particle emission from ²¹⁰Po. The hit points were determined based on coincidence of anode and cath- ode detections. Figure 5 (c) shows the anode-cathode plane for the track. The open circles are

Fig. 4: Energy spectrum for alpha particles from ²¹⁰Po (5*.*3 MeV).

 data. The red solid line is a linear fit result. The dashed line represents the edge of the sample win- dow. The solid blue point is the emission point of the alpha particle. The scheme of the determina- tion of the emission point, or the track sense, is explained in Sec. 3.4. Figure 5 (a) and (d) show anode- and cathode-drift planes, respectively. The drift coordinate is converted from the timing and is set to zero base, which corresponds to the drift- plate position. Figure 5 (b) shows a flash ADC waveform.

 The track angles were determined on the anode- cathode, anode-drift, and cathode-drift planes. These angles were determined with a common fit- ting algorithm. First, the weighted means of the ¹⁹² hit points (x_w, y_w) were defined as

$$
\left(\begin{array}{c} x_{\rm w} \\ y_{\rm w} \end{array}\right) = \frac{1}{n} \sum_{j=0}^{n} \left(\begin{array}{c} x_{j} \\ y_{j} \end{array}\right),\tag{1}
$$

¹⁹³ where x_j and y_j are the measured hit points and *n* $_{194}$ is a number of points. Next, the track was shifted $_{214}$ 195 and rotated through the angle θ as follows

$$
\begin{pmatrix} x'_j \\ y'_j \end{pmatrix} = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} x_j - x_w \\ y_j - y_w \end{pmatrix} . \tag{2}\n\begin{pmatrix} 246 \\ 22 \\ 218 \end{pmatrix}
$$

 $\lim_{j\to\infty}$ Here x'_j and y'_j are the points after the shift, and ¹⁹⁷ rotation and the angle *θ* were determined to mini-198 mize the quantity f , which is defined as

$$
f(\theta) = \sum y'^2_{j},\tag{3}
$$

 where this formula means a sum of the square of the distance between the rotated point and the *x* axis. 2 This method has the advantage to determining the 1 angle with no infinity pole at $\theta = 90^\circ$, in contrast with a sample linear fit. −15 −10 5 10 15 16 16 16 17 16 17 16 17 16 17 16 17 16 17 16 17 16 17 16 17 16 17 16 17 16 17 16 17 16 17 16 1

Fig. 5: Event display of an alpha particle from 210Po. (a) cathode-drift projection, (b) flash ADC waveform (c) cathode-anode projection, and (d) anode-drift projection are displayed. The drift coordinate is set to zero base corresponding to the drift plate position for the top of the track.

3.4. Track-sense determination

 Backgrounds in low radioactivity alpha-particle detectors are in general alpha particles from the $_{207}$ radon (radon- α) and material in the detector (detector-*α*). The radon-*α*s are expected to be dis- tributed uniformly in the gas volume with isotropic directions. The detector-*α*s are expected to have position and direction distributions specific to their sources. One of the main sources of the detector-*α*s $_{213}$ are the μ -PIC and the directions are mostly upward. Since the direction of alpha particles from the sample are downward, these detector-*α*s and half of the radon-*α*s can be rejected by the cut of upward-direction events.

 The deposit energy per unit path length, dE/dx of an alpha particle with an initial energy over a few MeV, has a peak before stopping (Bragg peak). The number of electrons ionized by the alpha parti- cle in the gas is proportional to dE/dx , and dE/dx along the track profile is projected onto the time ²²⁴ evolution in the signal due to the mechanism of the 225 TPC. This time profile was recorded as the wave-²²⁶ form and thus the track sense (i.e., whether the tack 1800 ²²⁷ was upward or downward) can be determined from

²²⁸ the waveform.

229 A parameter to determine the track sense is

$$
F_{\rm dwn} = S_2/(S_1 + S_2),\tag{4}
$$

230 where S_1 and S_2 are the time-integrated waveform 231 before and after the peak. They are defined as

$$
S_1 = \int_{t_0}^{t_p} v(t)dt,
$$
 (5)

$$
S_2 = \int_{t_p}^{t_1} v(t)dt.
$$
 (6)

232 Here, $t_0 = 0 \mu s$, $t_1 = 1.5 \mu s$, and t_p are the start, 233 stop, and peak time, respectively, for the waveform $_{234}$ shown in Fig. 5 (b). Figure 6 shows typical F_{dwn} ²³⁵ distribution with the alpha-particle source, where 1 ²³⁶ most of the events are expected to be downward- $_{237}$ oriented. The F_{dwn} values of the downward events ²³⁸ are distributed around 0.7, as shown by the black-²³⁹ shaded histograms. Conversely, radon-*α*s have an ²⁴⁰ isotropic direction, i.e., F_{dwn} has two peaks, as ²⁴¹ shown by the red solid histogram, where the radon- α are background events in the sample test data, ²⁴³ as explained later. The scale of the source- α was ²⁴⁴ normalized to the radon- α peak of downward for ²⁴⁵ clarity. The selection efficiency of $F_{\text{dwn}} > 0.5$ was ²⁴⁶ estimated to be 0.964 ± 0.004 in the source- α spec-²⁴⁷ trum while the radon background was reduced to 0 $_{248}$ half. The blue dashed histogram is a spectrum that 2 ²⁴⁹ subtracted the normalized source- α from the radon- α . The cut efficiency of the upward-oriented events $_{251}$ ($F_{\text{dwn}} \leq 0.5$) was estimated to be 0.85 ± 0.04 . The ₂₇₈ ²⁵² energy dependence of F_{dwn} will be explained in Sec. $_{279}$ ²⁵³ 3.6.

²⁵⁴ *3.5.* Distribution of emission position

²⁵⁵ Since alpha particles are mainly emitted from the ²⁸³ source, the top points of the alpha-particle tracks ²⁵⁷ trace the shape of the radioactivity on the sample. ²⁸⁵ Figures 7 (a) and 7 (b) show the anode–cathode projection distribution of the top and bottom of the alpha-particle tracks, respectively, where the top and bottom are defined as the zero and maximum drift coordinate, respectively, as shown in Figs. 5 $_{263}$ (a) and 5 (d). The dashed line represents the edge $_{291}$ of the drift-plate sample window. Comparing Fig. 7 (a) with Fig. 7 (b) clearly reveals the shape of the radioactivity.

Fig. 6: Downward-oriented distribution for source-*α* (black shade), radon- α (red solid), and a histogram made by subtracting the radon- α spectrum from the source- α one(blue dashed).

 The position resolution was evaluated along the four dashed lines in Fig. 7 (a). The number of events was projected onto the axis perpendicular to the lines and was fit with error functions. As a result, the position resolution was determined to be 0.68 ± 0.14 cm (σ) , where the error is a standard deviation in the four positions.

²⁷⁴ *3.6.* Efficiency of event selection

²⁷⁵ *3.7.* Detection and selection efficiency

²⁷⁶ To select good events for alpha particles from the ²⁷⁷ sample, we use the following criteria: (C1) selection for events with good fitting tracks, $(C2)$ cut for the upward-oriented events, and $(C3)$ selection ²⁸⁰ for events with emission points in the sample region. ²⁸¹ For criterion C1, the good fit to track events was selected as $f_{\min}(\theta)/(n-1) < 0.02$ cm² for the anode-cathode, anode-drift, and cathode-drift planes to remove events that had any noise and to remove candidates for electron tracks, where $f_{\text{min}}(\theta)$ is a minimum of Eq. (3).

Criterion C2 rejects the upward-oriented tracks ²⁸⁸ with $>$ 3.5 MeV and $F_{\text{dwn}} \leq 0.5$ because the de-
²⁸⁹ termination efficiency depends on the energy. The termination efficiency depends on the energy. The upward- and downward-oriented tracks can be determined with 95% or more certainly at over ²⁹² 3*.*5 MeV. Note that this cut was applied for the ²⁹³ events *>* 3*.*5 MeV, because the radon background, ²⁹⁴ which was assumed to be the dominant background

source. The dashed line is the edge of the sample window. Fig. 7: Anode–cathode projection distributions of (a) top and (b) bottom of tracks for alpha particles emitted from the

500 500 contribution to the energy range below 3*.*5 MeV ²⁹⁵ source, created the peak around 6 MeV and the was limited.

 – source- α , and thus the it was a negligible. cathode coordinate. The rate of radon- α in the se-200 cm region for arpin particle of mission point was see $\frac{1}{299}$ the remained the radon and detector- α s, the selec- $_{298}$ For criterion C3, as shown in Fig. 7 (a), to reject tion region for alpha-particle emission point was set lected region was less than a few hundred time of

305 The selection efficiency for C1, C2, and 336 C3 containing the detection efficiency was $_{307}$ (2.17 ± 0.29) \times 10⁻¹ counts/ α (the ratio of the 308 count rate to the α rate of the source), where 339 the error represents the systematic error of C1 to C3 selections and uncertainly of the source

radioactivity and the statistic error is negligible.

3.8. Sample test and background estimate

3.8.1. Setup

 $A \ 5 \times 5 \text{ cm}^2$ piece of the standard μ -PIC whose α rate was known to be 0.28 ± 0.12 *α*/cm²/hr in previous work [14] served as a sample and was in- spected by using the detector. The setup is shown in Fig. 8. The live time was 75.85 hr.

Fig. 8: Setup for a 5×5 cm² piece of the standard μ -PIC as sample.

3.8.2. Background in sample region

321 tracting the background rate. Considered back- $\frac{320}{20}$ The α rate of the sample was estimated by sub- ground was mainly the radon- α . The detector mea- sured both the α rates on the region of the sample and around the sample (outer region). The back- ground rate could be determined from the α rate in $\frac{326}{120}$ the outer region. The net α rate from the sample was thus evaluated by subtracting the background rate from the rate of the sample region. It was nec- essary to confirm that the background rates in both regions were consistent with each other.

³³⁴ Measured energy spectra are shown in Fig. 3. The
³³⁵ red- and black-shaded histograms show the energy 331 We checked the upward-oriented $(F_{\text{dwn}} \leq 0.5)$ α rate in both regions because the alpha particles from a sample are typically emitted downward. Measured energy spectra are shown in Fig. 9. The spectra inside and outside the sample region, respectively. These spectra are scaled by the selection efficiency. Both peaks are around 6 MeV and *α* rates are $(2.16^{+0.54}_{-0.35}) \times 10^{-2}$ (inside) and $(1.54^{+0.64}_{-0.40}) \times 10^{-2} \alpha/cm^2/hr$ (outside). Therefore, the background condition inside the sample region

³⁴² is consistent with the background condition outside ³⁴³ the sample region. The alpha-particle energy spec-³⁴⁴ trum is interpreted as the radon peaks at 5.5 MeV 372
373
374
375
376

³⁴⁵ (²²²Rn), 6.0 MeV (²¹⁸Po), and 7.7 MeV (²¹⁴Po).

 346 The downward-oriented $(F_{\text{dwn}} > 0.5)$ α rate out-346 The downward-oriented ($P_{\text{dwn}} > 0.5$) *α* Tate out-

347 side the sample is $(1.58^{+0.29}_{-0.26}) \times 10^{-2}$ α/cm²/hr, as 348 shown in the black-shaded spectrum of Fig. 11. In this work, the background rate was improved by one $\frac{1}{3}$ 350 order of magnitude in comparison with that of our $_{351}$ previous work [14]. The background reduction is at- $_{\rm 352}$ $\,$ tributed to the track-sense determination to reject ³⁵³ upward-oriented alpha (for > 3.5 MeV) and the re- $\frac{1}{354}$ placement of the low-*α* μ -PIC (for \leq 3.5 MeV). In ³⁵⁵ the energy region between 2.0 and 4.0 MeV, where $\frac{356}{356}$ most radon background is suppressed, the back- σ ₃₅₇ ground rate is $(9.6^{+7.9}_{-5.6}) \times 10^{-4} \alpha/cm^2/hr$. $\frac{1}{2}$ e downward-oriented $(r_{\text{dwn}} > 0.5)$ α rate outis $(9.6^{+7.9}_{-5.6}) \times 10^{-4} \alpha/cm^2/hr$. ⁰ ¹ ² ³ ⁴ ⁵ ⁶ ⁷ ⁸ ⁹ ¹⁰ ⁰ $\frac{\text{th}}{\text{tr}}$
 $\frac{(22)}{\text{SiC}}$
 $\frac{\text{SiC}}{\text{th}}$
 $\frac{\text{th}}{\text{c}}$ $_{349}$ thi ա) $\rm _{int}$ −10

Fig. 9: Downward-oriented alpha-particle energy spectra inside (red) and outside (black shade) the sample region.

³⁵⁸ *3.8.3. α* rate of sample 0.1

 Figure 10 shows the distribution of the top of the tracks for the sample, where the candidates are se-0.05 $_{361}$ lected by the criteria C1 and C2. The regions Ω $_{390}$ 362 and 2 are sample and background regions, respec- $\frac{363}{100}$ tively. The sample region is the inside of ± 5 cm of anode and cathode. The background region is the outside of the sample region and the inside of *±*7*.*5 cm of anode and cathode. Figure 11 shows the energy spectra of downward-oriented alpha par- ticles in the sample (red) and the background re- $\frac{1}{369}$ gion (black shaded). The α rate of the sample $\frac{1}{398}$ $\frac{\text{was calculated to be } (3.57_{-0.33}^{+0.35}) \times 10^{-1} \text{ }\alpha/\text{cm}^2/\text{hr}}{2.57_{-0.33}^{+0.35}}$ (*>* 2*.*0 MeV) by subtracting the background rate.

6 −5 379 is consistent with the prediction of prior measure-377 using the HPGe detector. Although the error is impurities of $\overline{1}$ in and $\overline{0}$ are measured to be $\frac{1}{374}$ be 6.0 ± 1.4 or 3.0 ± 0.7 ppm, respectively. The ³⁷² Assuming the alpha spectrum is constituted only $\overline{5}$ observe the background arphas at the same to −5 ³⁸⁰ ment. In this sample test, it was demonstrated to 378 huge because of the continuous energy spectrum, it 5.84 ± 0.03 and 2.31 ± 0.02 ppm, respectively, by $\frac{1}{375}$ impurities of $\frac{232 \text{ Th}}{375}$ and $\frac{238 \text{ U}}{380}$ are measured to be $\frac{1}{373}$ from $\frac{232 \text{Th}}{373}$ or $\frac{238 \text{U}}{373}$, the impurity is estimated to ³⁸¹ observe the background alphas at the same time.

particle track. The regions \hat{Q} and \hat{Q} are the sample and background regions, respectively. Fig. 10: Distribution of the top of downward-oriented alpha-

0.8 ³⁸² **4. Discussion**

391 statistical error was $\sigma \sim 3 \times 10^{-3}$. With a measure-2 2 ³⁹⁰ this work, the live time was only three days, and the $\frac{1}{289}$ of the measurement time (*t*) given as $\sigma \propto 1/\sqrt{t}$. In 10^{-2} *α*/cm²/hr. The statistical error (*σ*) is ex-8 ³⁸⁵ surements. In this energy range, the background is 10 8 ³⁸⁴ ergy between 2 and 9 MeV based on long-term mea-12 ³⁸³ We begin by discussing the sensitivity for the en-³⁹¹ Statistical error was $\frac{\partial}{\partial x} \times 10^{11}$. With a measure-
³⁹² ment time of one month, the error of radon-*αs* was 388 pected to scale with the inverse of the square root ergy between 2 and 9 MeV based on long-term measurements. In this energy range, the background is dominated by the radon-*αs* with $\sim (1.58^{+0.29}_{-0.26}) \times$ estimated to be $\sigma \sim 1 \times 10^{-3} \alpha/cm^2/hr$. When the ³⁹⁴ *α* rate as the same of the radon-*α*s was observed, the sum of squares of these σs for the sample and radon-³⁹⁶ *a*s would be expected to be a few $10^{-3} \alpha/cm^2/hr$ as the measurement limit by subtraction with these α rates.

³⁹⁹ The edges region (anode *∼ ±*15 cm or cathode ⁴⁰⁰ *∼ ±*15 cm) has a high rate of background, as shown

 \tilde{a} Fig. 11: Downward-oriented alpha-particle energy spectra in sample region (red) and background region (black shade).

 $_{411}$ a next improvement, a material with less radiative 0.05 ⁴⁰⁹ by the fiducial region cut, these impurities could α rate was (1.16 ± 0.06) \times 10⁻¹ *α*/cm²/hr. Al-0.15 ⁴⁰⁵ tion area. As an impurity candidate, a piece of 0.2 ⁴⁰³ ticles. The alpha particles were mainly oriented 0.25 ⁴⁰¹ in Fig. 10. These events have an energy and path-. 11
ple
Fig
tles.
wai
n a μ_{412} impurities should be used for the PCB. ⁴⁰² length dependence similar to that of the alpha par-⁴⁰⁴ upward and were emitted from outside the detec-⁴⁰⁶ the printed circuit board (PCB) was inspected and ⁴⁰⁸ though the alpha-particle events could be rejected ⁴¹⁰ be the radon sources (see Fig. 12). Therefore, as

Fig. 12: Schematic cross section of background alpha particles in detector setup.

⁴¹³ The goal for detector sensitivity is less than

 $\frac{1}{2}$. With these improvements, the detector wo $\,$ argon gas [18]. A recent NEWAGE detector sup-0.14 0.001 ⁴¹⁷ the cooled charcoal and using a material with less 0.2 15 15 ⁴¹⁶ potentially improve the background rate by using ¹¹/²¹/²¹¹/² 1^{-4} α /cm²/hr which corresponds to measuri 5 $\begin{bmatrix} \mathbf{r} \ \mathbf{r} \ \mathbf{p} \ \mathbf{t} \ \mathbf{u} \end{bmatrix}$ −5 -−5 ⁴²³ achieved to the goal of performance. $\frac{421}{2}$ presses the radon to $\frac{1}{50}$ by using cooled charcoal 5 ⁴¹⁹ charcoal could suppress the radon by 99% in the $_{418}$ impurities. A recent study reported that a cooled 422 [5]. With these improvements, the detector would h_{414} 10⁻⁴ *α*/cm²/hr, which corresponds to measuring

−15 ⁴²⁴ **5. Conclusion**

₄₂₅ We developed a new alpha-particle imaging de- \sim ⁴³⁵ another measurement. A measurement of the alpha 433 a piece of the standard μ -PIC was measured as a 1431 0.964 \pm 0.004 and the cut efficiency of the upward- $\frac{1}{2}$ $\frac{1}{2}$ ⁴²⁷ sured energy resolution is 6.7% (σ) for 5.3 MeV al-0 [−]²⁰ [−]¹⁵ [−]¹⁰ [−]⁵ ⁰ ⁵ ¹⁰ ¹⁵ ²⁰ [−]²⁰ [−]²⁰ [−]¹⁵ [−]¹⁰ [−]⁵ ⁰ ⁵ ¹⁰ ¹⁵ ²⁰ [−]²⁰ ⁴²⁶ tector based on the gaseous micro-TPC. The mea- $\frac{6}{438}$ near the radon-*α* $((1.58^{+0.51}_{-0.42}) \times 10^{-2} \text{ }\alpha/\text{cm}^2/\text{hr})$ 8 ⁴³⁷ established at the same time. A background rate ⁴³⁴ sample, and the result is consistent with the one by 432 oriented events is 0.85 ± 0.04 at > 3.5 MeV. Also, 16 ⁴³⁰ the downward-oriented events' selection efficiency is 428 pha particles. The measured position resolution 8 0.08 ⁴³⁶ particles from a sample and background was also $\frac{1}{429}$ is 0.68 ± 0.14 cm. Based on a waveform analysis, ⁴³⁹ was achieved.

2 0.02 ⁴⁴⁰ **Acknowledgments**

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⁴⁵⁰ **References**

- ⁴⁵¹ [1] R Bernabei, et al., J. Phys. Conf. Ser. **1056** (2018) 452 012005.
453 [2] XENOI
- ⁴⁵³ [2] XENON Collaboration, Eur. Phys. J. **77** 881 (2017).
- ⁴⁵⁴ [3] D. S. Akerib, et al., Phys. Rev. Lett. **118** 021303 (2017).
- ⁴⁵⁵ [4] T. Tanimori, et al., Phys. Lett. B **578** (2004) 241.
- ⁴⁵⁶ [5] K. Nakamura, et al., Prog. Theo. Exp. Phys. (2015) $457 \overline{)043F01}$
- ⁴⁵⁸ [6] The GERDA Collaboration, Nature **544** (2017) 47.
- ⁴⁵⁹ [7] K.Asakura, et al., Nucl. Phys. A **946** (2016) 171.
- ⁴⁶⁰ [8] R. Arnold, et al., Eur. Phys. J. C **78** (2018) 821.
- [9] R. Arnold, et al., PRL **119**, 041801 (2017).
- [10] A. S. Barabash, et al., JINST **12** (2017) P06002.
- [11] K. Abe, et al., Nucl. Instr. Meth. A **884** (2018) 157.
- [12] K. Miuchi, et al., Phys. Lett. B **686** (2010).
- [13] K. Miuchi, et al., Phys. Lett. B **654** (2007) 58.
- [14] T. Hashimoto, et al., AIP Conf. Proc. **1921**, 070001 (2018).
- [15] T. Hashimoto, et al., in preparation.
- [16] R. Orito, et al., IEEE Trans. Nucl. Scie. **51**, 4 (2004) 1337.
- [17] H. Kubo, et al., Nucl. Instr. Meth. A **513** (2003) 93.
- [18] M. Ikeda, et al., Radioisotopes, **59**, (2010) 29.