Elsevier Editorial System(tm) for Nuclear Inst. and Methods in Physics Research, A Manuscript Draft

Manuscript Number:

Title: Performance check of the CsI(Tl) calorimeter for the J-PARC E36 experiment by observing  $e^+$  from muon decay

Article Type: Full length article

Section/Category: High Energy and Nuclear Physics Detectors

Keywords: Kaon decay; CsI(Tl) Calorimeter; Waveform Analysis

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# Performance check of the CsI(Tl) calorimeter for the J-PARC E36 experiment by observing $e^+$ from muon decay

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## Abstract

The J-PARC E36 experiment is searching for a lepton universality violation with a stopped kaon beam by measuring the ratio of the  $K^+$  decay widths  $\Gamma(K_{e2})/\Gamma(K_{\mu 2}) = \Gamma(K^+ \to e^+\nu_e)/\Gamma(K^+ \to \mu^+\nu_{\mu})$ . Since the radiative  $K^+ \to e^+\nu_e\gamma$  decays are backgrounds to be removed in this measurement, the radiated  $\gamma$  rays were detected in a CsI(Tl) calorimeter. The energy calibration for the 768 CsI(Tl) modules was performed using mono-chromatic  $\mu^+$ s from the  $K_{\mu 2}$  decays. The delayed  $e^+$  signals from the muon decays were required in order to improve the S/N ratio of the  $K_{\mu 2}$  peak by suppressing background events. In addition, a new energy calibration method of the CsI(Tl) calorimeter using stopped cosmic muons has been established.

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Keywords: Kaon decay, CsI(Tl) Calorimeter, Waveform Analysis

## 1 1. Introduction

The  $K^+ \to l^+ \nu_l$  decay channel is one of the best processes to search for a lepton universality violation [1-3]. The ratio of  $K^+ \to e^+ \nu_e$  ( $K_{e2}$ ) and  $K^+ \to \mu^+ \nu_\mu$  ( $K_{\mu2}$ ) decay widths ( $R_K$ ) can be very precisely calculated in the framework of the Standard Model (SM) under the assumption of  $\mu$ -e universality as [4],

$$R_K^{\rm SM} = \frac{\Gamma(K_{e2})}{\Gamma(K_{\mu 2})} = (2.477 \pm 0.001) \times 10^{-5}.$$
 (1)

9 In order to compare the experimental  $R_K$ 10 value with the SM prediction, the internal bremsstrahlung process in radiative  $K^+ \to e^+ \nu_e \gamma$  $(K_{e2\gamma}^{\text{IB}})$  and  $K^+ \to \mu^+ \nu_\mu \gamma$   $(K_{\mu2\gamma}^{\text{IB}})$  decay has to be included in the  $K_{e2}$  and  $K_{\mu2}$  samples. On the other hand, the structure dependent processes in radiative  $K^+ \to e^+ \nu_e \gamma$   $(K_{e2\gamma}^{\text{SD}})$  and  $K^+ \to \mu^+ \nu_\mu \gamma$   $(K_{\mu2\gamma}^{\text{SD}})$ decays are backgrounds and should be removed in the analysis [3]. A deviation of the experimentally measured  $R_K$  from the SM value would lead to a  $\mu^$ e universality violation and indicate the existence of New Physics beyond the SM.

The J-PARC E36 experiment aims to make a precise  $R_K$  measurement by adopting a stopped  $K^+$ beam method [5, 6]. A separated 800 MeV/c  $K^+$ beam was slowed down by a degrader and stopped in a position sensitive  $K^+$  stopper. The momentum measurement of the charged particles was performed using a 12-sector ion-core superconducting toroidal spectrometer, as shown in Fig. 1. The ra-

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Preprint submitted to Nucl. Instr. Meth. A

diated photon from the above radiative processes 29

was measured by a CsI(Tl) calorimeter, an assem-30

bly of 768 CsI(Tl) crystals, which covers 75% of the 31

total solid angle. The photon energy and hit posi-32 tion were obtained by summing the energy deposits 33 and by determining the energy-weighted centroid, 34 respectively. Since the SD component subtraction 35 is one of the key issues in E36, the understanding 36 of the CsI(Tl) performance is very important. 37

This paper is organized as follows. Details of the 38 CsI(Tl) calorimeter and the analysis procedure are 39 described in Section 2 and Section 3. In Section 4, a 40 calibration method using the mono-chromatic  $\mu^+$ s 41 from the  $K_{\mu 2}$  decays is explained. A new method 42 of the CsI(Tl) energy calibration using stopped 43 cosmic-ray muons is discussed in Section 5. The 44 results obtained in the present studies are summa-45 rized in Section 6. 46

#### 2. CsI(Tl) calorimeter 47

The CsI(Tl) calorimeter was originally con-48 structed for the KEK-PS E246 experiment to search 49 for a T-violating transverse muon polarization in 50  $K^+ \to \pi^0 \mu^+ \nu_\mu$  decay [7–9]. There are 12 holes for 51 outgoing charged particles and 2 holes for the beam 52 entrance and exit. Each crystal had a coverage of 53  $7.5^{\circ}$  along both the polar and azimuthal directions. 54 The length of the CsI(Tl) crystal was 25 cm which 55 was enough to obtain sufficient energy resolution as 56 well as avoid nuclear counter effects. 57

Since the CsI(Tl) calorimeter had to be oper-58 ated under a relatively strong fringing field from 59 the toroidal magnet where PMTs would be diffi-60 cult to use, PIN photodiodes (PIN diodes) were 61 employed to read out the scintillation light of the 62 CsI(Tl) crystals. Each crystal with its associated 63 PIN diode and pre-amplifier was assembled in an 64 Al container of 0.1 mm thickness. A charge sen-65 sitive pre-amplifier with a time constant of 600  $\mu$ s 66 and a gain of 0.5 V/pC was attached directly to 67 the PIN diode. The output signal from the pre-68 amplifier was fed to a shaping amplifier with 1  $\mu$ s 69 shaping time. The waveforms of the shaping am-70 plifier outputs were recorded by VF48 flash ADC 71 manufactured by the TRIUMF national laboratory 72 100 [10]. The VF48 had a 10  $\mu$ s time range and was op-73 erated with a 25 MHz external clock signal (1 TDC 74 channel = 40 ns). 75

## 3. Waveform analysis

## 3.1. Waveform model

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The  $\gamma$ -ray energy and timing can be determined by fitting the CsI(Tl) output signal using a dedicated waveform model function. A typical waveform from the CsI(Tl) calorimeter is shown in the Fig. 2 (a), as indicated by the black open circles. In the analysis, we adopted the following waveform formula as,

$$f(t) = \frac{A}{1 - \exp\{-(t - \tau_0)/\lambda\}} \cdot Freq\left(\frac{t - \tau_0 - d}{\mu}\right)$$
$$\left\{\frac{t - \tau_0}{\tau_1} \exp\left(1 - \frac{t - \tau_0}{\tau_1}\right)$$
$$+ \varepsilon \frac{t - \tau_0}{\tau_2} \exp\left(1 - \frac{t - \tau_0}{\tau_2}\right)\right\},$$
(2)

where A is the amplitude of the pulse and  $\tau_0$  is the rise time. The  $\lambda, \mu$  and  $\tau_1, \tau_2$  parameters are time constants to express the rise and decay parts of the pulse, respectively.  $d \sim 1 \,\mu s$  is introduced for a timing adjustment and  $\varepsilon \sim 0.06$  is the ratio of the two decay components. Freq(x) is known as the frequency function given as

$$Freq(x) = \frac{1}{\sqrt{2}} \int_{\infty}^{x} exp(-t^2/2)dt.$$
 (3)

Then, in order to determine these parameters, an equal-weighted  $\chi^2$  quantity is introduced as,

$$\chi^2 = \sum_{i=1}^{250} \left\{ A_i - f(t_i) \right\}^2, \tag{4}$$

where  $A_i$  and  $t_i$  are the ADC value and time of the *i*th waveform points, respectively. The parameters in the model function were derived by minimizing the  $\chi^2$  values. The red line in Fig. 2 (a) is the fitting result using the above method, and the deviation of each data point (dh) is shown in fig. 2 (b). Typical  $\chi^2$  values are distributed in the region of 100-500 which is mainly due to the imperfect reproducibility of the CsI(Tl) output by the waveform model.

## 3.2. Pulse separation of pileup events

For the analysis of pileup events, the maximum dh value  $(dh_{\rm max})$  was first determined in the entire region using a single-pulse fitting. The waveforms with  $|dh_{\rm max}| > 10$  can be recognized as two or more than two pulse components. The deviation of the data points from the single-pulse fitting result for a typical pileup event is shown in Fig. 3 (b) as



Fig. 1: Cross sectional end and side views of the setup for the J-PARC E36 experiment. The momentum vectors of charged particles and photons are determined by the toroidal spectrometer and the CsI(Tl) calorimeter, respectively.



Fig. 2: (a) Typical waveform of the CsI(Tl) calorimeter signal. The open circles are the data and the red line is a fitting result of the waveform model. (b) The deviation of the data points from the fitting result.

the black line. These events were treated as pileup 104 events, and multiple pulses in the fitting were taken 105 into account. Then, the  $\chi^2$  value using a double-106 pluse waveform was again minimized by changing 107 the fitting parameters. A typical pileup waveform 108 is shown in Fig. 3 (a) as the black open circles. We 109 can accept events as a double-pulse waveform with 110 the conditions of (i) a waveform with  $|dh_{\text{max}}| < 10$ , 111 (ii) the time interval between the 1st and the 2nd 112 signals is longer than 200 ns. The rejected events 113 are treated as events with further multiple signals. 114 The red and green solid lines in Fig. 3 (a) are the fit-115 ting results using the single-pulse and double-pulse 116 fitting functions, respectively. The associated de-117

composed pulses are shown as the green (1st pulse) and blue (2nd pulse) dotted lines. The red line in Fig. 3 (b) shows the *dh* distribution assuming the double-pulse fittings, which indicates successful pulse separation using the double-pulse fitting.



Fig. 3: (a) Typical pileup waveform of the CsI(Tl) calorimeter signal. The open circles are the data points. The red and green lines are the results adopting the single- and doublepulse fitting function. The green and blue dotted lines are the decomposed 1st and 2nd pulses. (b) The deviation of each data point from the fit curves. The black and red lines are the results using the single and double fitting.

#### 4. CsI(Tl)calibration using $K_{\mu 2}$ decay 123 events 124

4.1. Background reduction by observing the  $e^+$ 125 from  $\mu^+$  decay 126

The CsI(Tl) energy calibration was performed us-127 ing mono-chromatic  $\mu^+$ s from the  $K_{\mu 2}$  decays at 128 rest in the  $K^+$  stopping target. The  $\mu^+$  kinetic 129 energy was 152.5 MeV. These muons were stopped 130 in the CsI(Tl) crystal and generated the delayed 131  $e^+$  signal from the subsequent  $\mu^+ \to e^+ \bar{\nu_{\mu}} \nu_e$  decay. 132 The  $e^+$  signal can be observed as the second pulse 133 in the waveform analysis using the double-pulse fit-134 135 ting.

The  $K_{\mu 2}$  events were selected by the following 136 conditions: (I) the number of hit crystals was only 137 one, (II) the first pulse time coincided with the  $K^+$ 138 decay, and (III) the waveform data was successfully 139 analyzed as a double-pulse waveform. 140

The pulse height spectrum obtained by selecting 141 events with the conditions of only (I) (II) are shown 142 in Fig. 4 as the black histogram. On the other 143 hand, the red histogram represents events selected 144 with all the above conditions. It is clearly seen that 145 background components below the  $K_{\mu 2}$  peak are 146 significantly removed by requiring the  $\mu^+$  decay in 147 the CsI(Tl). Here, the backgrounds are considered 148 to be mainly accidental events created by the beam 149 particles. 150

Then, the signal to noise ratio (S/N) was calcu-151 lated as, 152

$$S/N = \frac{N(500 \le l < 800)}{N(l < 500, \ 800 \le l)},\tag{5}$$

where l is the pulse height of the first pulse ob-153 tained by the fitting. The  $K_{\mu 2}$  peak region and 154 the background dominant region were separated as 155  $N(500 \le l \le 800)$  and  $N(l \le 500, 800 \le l)$ , respec-156 tively. The S/N ratio was determined to be 0.42157 for the events selected with the conditions of (1)(2). 158 Next the  $\mu^+$  selection by requiring the double-pulse 159 waveform was performed, and the S/N was ob-160 tained to be 4.42. Thus, we can conclude that the 161 requirement of the  $\mu^+$  stop and decay in the CsI(Tl) 162 is a very useful technique to reduce the backgrounds 163 from the beam particles and make the CsI(Tl) en-164 193 ergy calibration significantly more accurate. 165

#### 4.2. CsI(Tl) performance check 166

For the CsI(Tl) energy calibration, the  $\mu^+$  en- 196 167 ergy loss in the target system should be added to 197 168



Fig. 4: Integrated pulse-height spectrum. The black spectrum shows the events selected with the conditions of (1)(2). The red shaded histogram shows the events selected with all the conditions. The region indicated by the two dotted lines is used to estimate the S/N ratio.

the  $\mu^+$  energy observed by the CsI(Tl). The energy conversion factor, k, can be formulated as  $k = (152.5 - E_t \text{ MeV})/l$ , where  $E_t$  is the muon energy loss in the target. The  $\mu^+$  path length in the target was obtained by connecting the CsI(Tl) center of the  $\mu^+$  hit module and the  $K^+$  vertex position determined by the target system. The typical k value was obtained to be 2.1–2.5 MeV<sup>-1</sup>. Then, the  $\mu^+$  energy spectrum from the  $K_{\mu 2}$  decay is obtained by taking into account the energy loss in the target as  $E = kl + E_t$ , as shown in Fig. 5. The red and blue spectra indicate the calibrated energy spectrum with and without the target energy correction, respectively. The target energy correction improved the energy resolution to  $\sigma=2.63\%$  from 4.73%.

Also, the CsI(Tl) timing information was checked by requiring the  $e^+$  signals to reduce the effects from accidental backgrounds. The 40 ns clock timing uncertainty of VF48 was corrected for by measuring the trigger signal in the same VF48 module. The timing resolution was determined to be  $10.7 \pm 0.1 \text{ ns} (\sigma).$ 

## 5. A new method of energy calibration using stopped comic-ray muons

It is possible to consider a new CsI(Tl) calibration method using stopped cosmic-ray muons with the subsequent  $e^+$  emission in the CsI(Tl) calorimeter. This method is proposed to measure the  $e^+$ 

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Fig. 5: The calibrated energy spectra obtained using the  $\rightarrow \mu^+ \nu_\mu$  decays. The red spectrum includes a correction  $K^{\cdot}$ for the energy loss in the target. The red lines are the fitting results assuming a Gaussian function.

energy spectrum for a rough CsI(Tl) energy cali-198 bration without the need for a radioactive beam. 199 Since the maximum  $e^+$  energy from the muon de-200 cay is 52.32 MeV, the energy calibration can be 201 performed by measuring the  $e^+$  energy after the 202 cosmic-ray muon stops in the CsI(Tl) crystal. The 203 233 cosmic muons stop homogeneously in the CsI(Tl), 204 234 and we do not need to consider the specific struc-205 ture of the CsI(Tl) calorimeter. 206

The energy distribution of the decomposed sec-207 ond pulse is shown in Fig. 6 as indicated by the 208 black dots. Here the calibration parameters ob-209 239 tained from the  $K_{\mu 2}$  decays were used. The blue 210 240 squares and the red open circles are the calculated 211  $e^+$  and  $e^-$  energy distribution, respectively, using 212 242 a Monte Carlo simulation based on a GEANT4 213 Electromagnetic shower leakage from the code. 214 muon stopped module was taken into account. The 215 red cross-hatched area and the green filled area are 216  $e^+$  and  $e^-$  energy distributions assuming the muon 217 247 yield ratio  $F_+/F_- = 1.1$  and 1.6, respectively [11– 218 15]. The energy resolution of 2.63% ( $\sigma$ ) obtained 219 from the  $K_{\mu 2}$  calibration result has been used. 220

In order to determinate the energy calibration pa- 249 221 rameters using stopped cosmic-ray muons, a com- 250 222 mon gain parameter relative to the energy coeffi- 251 223 cients obtained from the  $K_{\mu 2}$  calibration results was <sup>252</sup> introduced. The reduced  $\chi^2_{\nu}/\text{NDF}$  determined by <sup>253</sup> 224 225 comparing the experimental data with the simula- <sup>254</sup> 226 tion was calculated as a function of the the above 255 227 relative gain coefficient, as shown in Fig. 7, where 256 228 NDF is the number of degrees of freedom. The 257 229 black dots and open squares correspond to the re-258 230



Fig. 6: Energy distributions of  $e^+$  ( $e^-$ ) from stopped cosmic muons. An electromagnetic gamma shower was taken into account in the simulation. The black hatched area is not used in the fitting because of the online threshold setting of 20 MeV.

sults obtained by assuming  $F_+/F_- = 1.1$  and 1.6, respectively. It should be noted that the fitting region was chosen to be 20-60 MeV because the online energy threshold was set to 20 MeV. The black hatched area was not used for the fitting. Scattering of the  $\chi^2_{\nu}$  values is due to random smearing to account for the CsI(Tl) energy resolution The lines in the figure represent the fitting results using a parabolic function. As a result, the relative coefficients for  $F_+/F_- = 1.1$  and 1.6 were determined to be  $0.986 \pm 0.033$  and  $1.001 \pm 0.032$ , which indicates the gain coefficients obtained from the stopped cosmic muons are consistent with those from the  $K_{\mu 2}$ events at the 3–4% level. Therefore, the experimental data were in good agreement with the above two simulation models, indicating a correct understanding of the  $e^+$  and  $e^-$  behavior generated from the stopped muons.

The muon lifetime curve was also measured using the time interval between the 1st and 2nd pulses, as shown in Fig. 8 by the black dots. Fitting the data with an exponential function, the decay constant was determined to be  $2.06 \pm 0.03 \ \mu s$  $(\chi^2_{\nu}/NDF = 69.5/43)$ , as shown by the red line. The observed time constant is a little shorter than the theoretical value which indicates that most of the  $\mu^-$  events are captured by CsI nuclei and do not contribute to the above lifetime measurement.

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Fig. 7: Reduced  $\chi^2_{\nu}$  obtained by changing the relative gain coefficient. The black dots and open squares correspond to the results obtained by assuming  $F_+/F_- = 1.1$  and 1.6, respectively. The lines in the figure represent the fitting results using a parabolic function. The gain coefficients obtained from the stopped cosmic muons are consistent with those from the  $K_{\mu 2}$  events within 3–4%.

#### 6. Conclusion 259

A model function for the waveform analysis of 290 260 CsI(Tl) calorimeter in the E36 experiment has been 291 261 developed, and the information of the decomposed 262 293 second pulses can be used for the event selection. 294 263 The CsI(Tl) energy calibration was successfully 295 264 performed by choosing the  $K_{\mu 2}$  events and impos-265 ing the existence of the second pulses, and the S/N 266 ratio was significantly improved. Then, the CsI(Tl) 267 performance was carefully checked by studying the 300 268 energy and timing resolutions. 269

A new energy calibration method using stopped 270 303 cosmic muons is proposed. The energy and timing 271 304 of the delayed  $e^+$  ( $e^-$ ) signals were determined by 305 272 the decomposed second pulse in the double-pulse 273 307 waveform analysis. The observed energy spectrum 274 308 is consistent with the simulation calculation with 309 275 an accuracy of 3-4%, indicating the establishment 310 276 of a new calibration method without using any ac-277 celerator facilities. 278

#### Acknowledgement 279

This work was supported by a Grant-in-Aid for 280 Scientific Research (C), No. 15K05113, from the 281 Japan Society for the Promotion of Science (JSPS) 282 in Japan and by NSERC and NRC (TRIUMF) in 283 Canada. The authors thank H. Yamazaki for en-284 couragement in executing this work. We would like 285 to thank the J-PARC staff for the excellent beam 286 delivery during our experimental beamtime. 287



Fig. 8: Time interval between cosmic muons and the delayed  $e^+$   $(e^-)$  signals through the  $\mu \to e \nu \bar{\nu}$  decays. The black dots and the red line are the data and fitting function, respectively.

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## Side View

















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