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Performance check of the CsI(Tl) calorimeter for the J-PARC E36 experiment by observing *e* ⁺ from muon decay

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Abstract

The J-PARC E36 experiment is searching for a lepton universality violation with a stopped kaon beam by measuring the ratio of the K^+ decay widths $\Gamma(K_{e2})/\Gamma(K_{\mu2}) = \Gamma(K^+ \to e^+ \nu_e)/\Gamma(K^+ \to \mu^+ \nu_\mu)$. Since the radiative $K^+ \to e^+ \nu_e \gamma$ decays are backgrounds to be removed in this measurement, the radiated γ rays were detected in a CsI(Tl) calorimeter. The energy calibration for the 768 CsI(Tl) modules was performed using mono-chromatic μ^+ s from the K_{μ^2} decays. The delayed e^+ signals from the muon decays were required in order to improve the S/N ratio of the $K_{\mu2}$ peak by suppressing background events. In addition, a new energy calibration method of the CsI(Tl) calorimeter using stopped cosmic muons has been established.

Keywords: Kaon decay, CsI(Tl) Calorimeter, Waveform Analysis

¹ **1. Introduction**

The $K^+ \to l^+ \nu_l$ decay channel is one of the best ³ processes to search for a lepton universality vio-4 lation [1–3]. The ratio of $K^+ \to e^+ \nu_e$ (K_{e2}) and *K*⁺ $\rightarrow \mu^+ \nu_\mu$ (*K*_{μ 2}) decay widths (*R_K*) can be very ⁶ precisely calculated in the framework of the Standard Model (SM) under the assumption of μ –*e* universality as $[4]$,

$$
R_K^{\rm SM} = \frac{\Gamma(K_{e2})}{\Gamma(K_{\mu 2})} = (2.477 \pm 0.001) \times 10^{-5}.
$$
 (1)

⁹ In order to compare the experimental *R^K* ¹⁰ value with the SM prediction, the internal

bremsstrahlung process in radiative $K^+ \rightarrow e^+ \nu_e \gamma$ ¹² ($K_{e2\gamma}^{\text{IB}}$) and $K^+ \rightarrow \mu^+ \nu_\mu \gamma$ ($K_{\mu 2\gamma}^{\text{IB}}$) decay has to be 13 included in the K_{e2} and $K_{\mu2}$ samples. On the other ¹⁴ hand, the structure dependent processes in radiative $K^+ \to e^+ \nu_e \gamma$ ($K_{e2\gamma}^{\text{SD}}$) and $K^+ \to \mu^+ \nu_\mu \gamma$ ($K_{\mu 2\gamma}^{\text{SD}}$) ¹⁶ decays are backgrounds and should be removed in ¹⁷ the analysis [3]. A deviation of the experimentally 18 measured R_K from the SM value would lead to a μ – ¹⁹ *e* universality violation and indicate the existence ²⁰ of New Physics beyond the SM.

 The J-PARC E36 experiment aims to make a prez cise R_K measurement by adopting a stopped K^+ ²³ beam method [5, 6]. A separated 800 MeV/ $c K^+$ beam was slowed down by a degrader and stopped ∞ in a position sensitive K^+ stopper. The momen- tum measurement of the charged particles was per- formed using a 12-sector ion-core superconducting toroidal spectrometer, as shown in Fig. 1. The ra-

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²⁹ diated photon from the above radiative processes

³⁰ was measured by a CsI(Tl) calorimeter, an assem-

³¹ bly of 768 CsI(Tl) crystals, which covers 75% of the

³² total solid angle. The photon energy and hit posi-

 tion were obtained by summing the energy deposits and by determining the energy-weighted centroid, respectively. Since the SD component subtraction is one of the key issues in E36, the understanding of the CsI(Tl) performance is very important.

 This paper is organized as follows. Details of the CsI(Tl) calorimeter and the analysis procedure are described in Section 2 and Section 3. In Section 4, a a calibration method using the mono-chromatic μ^+ s ⁴² from the $K_{\mu2}$ decays is explained. A new method of the CsI(Tl) energy calibration using stopped cosmic-ray muons is discussed in Section 5. The results obtained in the present studies are summa-rized in Section 6.

⁴⁷ **2. CsI(Tl) calorimeter**

 The CsI(Tl) calorimeter was originally con- structed for the KEK-PS E246 experiment to search for a T-violating transverse muon polarization in $K^+ \to \pi^0 \mu^+ \nu_\mu$ decay [7–9]. There are 12 holes for outgoing charged particles and 2 holes for the beam entrance and exit. Each crystal had a coverage of 54 7.5[°] along both the polar and azimuthal directions. The length of the CsI(Tl) crystal was 25 cm which was enough to obtain sufficient energy resolution as well as avoid nuclear counter effects.

 Since the CsI(Tl) calorimeter had to be oper- ated under a relatively strong fringing field from the toroidal magnet where PMTs would be diffi- cult to use, PIN photodiodes (PIN diodes) were employed to read out the scintillation light of the CsI(Tl) crystals. Each crystal with its associated PIN diode and pre-amplifier was assembled in an Al container of 0.1 mm thickness. A charge sen- sitive pre-amplifier with a time constant of 600 *µ*s σ and a gain of 0.5 V/pC was attached directly to the PIN diode. The output signal from the pre-69 amplifier was fed to a shaping amplifier with $1 \mu s$ shaping time. The waveforms of the shaping am- plifier outputs were recorded by VF48 flash ADC manufactured by the TRIUMF national laboratory [10]. The VF48 had a 10 μ s time range and was op- erated with a 25 MHz external clock signal (1 TDC τ ⁵ channel = 40 ns).

⁷⁶ **3. Waveform analysis**

⁷⁷ *3.1. Waveform model*

The *γ*-ray energy and timing can be determined by fitting the CsI(Tl) output signal using a dedicated waveform model function. A typical waveform from the CsI(Tl) calorimeter is shown in the Fig. 2 (a), as indicated by the black open circles. In the analysis, we adopted the following waveform formula as,

$$
f(t) = \frac{A}{1 - \exp\{-(t - \tau_0)/\lambda)\}} \cdot Freq\left(\frac{t - \tau_0 - d}{\mu}\right) \cdot \frac{\left\{\frac{t - \tau_0}{\tau_1} \exp\left(1 - \frac{t - \tau_0}{\tau_1}\right)\right\}}{\tau_1} + \varepsilon \frac{t - \tau_0}{\tau_2} \exp\left(1 - \frac{t - \tau_0}{\tau_2}\right)\},
$$
(2)

⁷⁸ where *A* is the amplitude of the pulse and τ_0 is ⁷⁹ the rise time. The $λ, μ$ and $τ_1, τ_2$ parameters are ⁸⁰ time constants to express the rise and decay parts 81 of the pulse, respectively. $d \sim 1 \mu s$ is introduced ⁸² for a timing adjustment and *ε ∼* 0*.*06 is the ratio 83 of the two decay components. $Freq(x)$ is known as ⁸⁴ the frequency function given as

$$
Freq(x) = \frac{1}{\sqrt{2}} \int_{\infty}^{x} exp(-t^2/2)dt.
$$
 (3)

⁸⁵ Then, in order to determine these parameters, an $_{86}$ equal-weighted χ^2 quantity is introduced as,

$$
\chi^2 = \sum_{i=1}^{250} \left\{ A_i - f(t_i) \right\}^2,\tag{4}
$$

 \mathbf{B} ^{*si*} where A_i and t_i are the ADC value and time of the *i*th waveform points, respectively. The parameters in the model function were derived by minimizing ⁹⁰ the χ^2 values. The red line in Fig. 2 (a) is the fitting result using the above method, and the deviation of each data point (*dh*) is shown in fig. 2 (b). Typical ⁹³ χ^2 values are distributed in the region of 100−500 which is mainly due to the imperfect reproducibility of the CsI(Tl) output by the waveform model.

⁹⁶ *3.2. Pulse separation of pileup events*

 For the analysis of pileup events, the maximum *dh* value (*dh*max) was first determined in the en- tire region using a single-pulse fitting. The wave-100 forms with $|dh_{\text{max}}| > 10$ can be recognized as two or more than two pulse components. The deviation of the data points from the single-pulse fitting result for a typical pileup event is shown in Fig. 3 (b) as

Fig. 1: Cross sectional end and side views of the setup for the J-PARC E36 experiment. The momentum vectors of charged particles and photons are determined by the toroidal spectrometer and the CsI(Tl) calorimeter, respectively.

points from the fitting result. nal. The open circles are the data and the red line is a fitting Fig. 2: (a) Typical waveform of the CsI(Tl) calorimeter sigresult of the waveform model. (b) The deviation of the data

 the black line. These events were treated as pileup events, and multiple pulses in the fitting were taken ¹⁰⁶ into account. Then, the χ^2 value using a double- pluse waveform was again minimized by changing the fitting parameters. A typical pileup waveform is shown in Fig. 3 (a) as the black open circles. We can accept events as a double-pulse waveform with 111 the conditions of (i) a waveform with $|dh_{\text{max}}| < 10$, (ii) the time interval between the 1st and the 2nd signals is longer than 200 ns. The rejected events are treated as events with further multiple signals. The red and green solid lines in Fig. 3 (a) are the fit- ting results using the single-pulse and double-pulse fitting functions, respectively. The associated de composed pulses are shown as the green (1st pulse) 1000 and blue (2nd pulse) dotted lines. The red line in Fig. 3 (b) shows the *dh* distribution assuming the double-pulse fittings, which indicates successful 800 $_{122}$ pulse separation using the double-pulse fitting.

are the results using the single and double fitting. pulse fitting function. The green and blue dotted lines are Fig. 3: (a) Typical pileup waveform of the CsI(Tl) calorimeter signal. The open circles are the data points. The red and green lines are the results adopting the single- and doublethe decomposed 1st and 2nd pulses. (b) The deviation of each data point from the fit curves. The black and red lines

123 **4.** CsI(Tl) calibration using $K_{\mu2}$ decay ¹²⁴ **events**

 μ ₁₂₅ 4.1. Background reduction by observing the e^+ $from \mu^+ \ decay$

 127 The CsI(Tl) energy calibration was performed using mono-chromatic μ^+ s from the K_{μ^2} decays at rest in the K^+ stopping target. The μ^+ kinetic ¹³⁰ energy was 152.5 MeV. These muons were stopped ¹³¹ in the CsI(Tl) crystal and generated the delayed ¹³² e^+ signal from the subsequent $\mu^+ \to e^+ \bar{\nu_\mu} \nu_e$ decay. T_{133} The e^+ signal can be observed as the second pulse ¹³⁴ in the waveform analysis using the double-pulse fit-¹³⁵ ting.

136 The $K_{\mu2}$ events were selected by the following ¹³⁷ conditions: (I) the number of hit crystals was only ¹³⁸ one, (II) the first pulse time coincided with the K^+ ¹³⁹ decay, and (III) the waveform data was successfully ¹⁴⁰ analyzed as a double-pulse waveform.

¹⁴¹ The pulse height spectrum obtained by selecting $_{142}$ events with the conditions of only (I) (II) are shown ¹⁴³ in Fig. 4 as the black histogram. On the other ¹⁴⁴ hand, the red histogram represents events selected ¹⁴⁵ with all the above conditions. It is clearly seen that 146 background components below the $K_{\mu 2}$ peak are $_{147}$ significantly removed by requiring the μ^+ decay in ¹⁴⁸ the CsI(Tl). Here, the backgrounds are considered ¹⁴⁹ to be mainly accidental events created by the beam ¹⁵⁰ particles.

 $T₁₅₁$ Then, the signal to noise ratio (S/N) was calcu-¹⁵² lated as,

$$
S/N = \frac{N(500 \le l < 800)}{N(l < 500, 800 \le l)},\tag{5}
$$

 where *l* is the pulse height of the first pulse ob-154 tained by the fitting. The $K_{\mu2}$ peak region and the background dominant region were separated as $N(500 \le l \le 800)$ and $N(l \le 500, 800 \le l)$, respec- tively. The S/N ratio was determined to be 0.42 158 for the events selected with the conditions of $(1)(2)$. $\frac{159}{159}$ Next the μ^+ selection by requiring the double-pulse waveform was performed, and the *S/N* was ob- tained to be 4.42. Thus, we can conclude that the requirement of the μ^+ stop and decay in the CsI(Tl) is a very useful technique to reduce the backgrounds from the beam particles and make the CsI(Tl) en-ergy calibration significantly more accurate.

¹⁶⁶ *4.2. CsI(Tl) performance check*

For the CsI(Tl) energy calibration, the μ^+ en-

Fig. 4: Integrated pulse-height spectrum. The black spectrum shows the events selected with the conditions of (1)(2). The red shaded histogram shows the events selected with all the conditions. The region indicated by the two dotted lines is used to estimate the S/N ratio.

168 are not be explicion to the following the control plate in the system show in the system space of the expression of the express the μ^+ energy observed by the CsI(Tl). The en-¹⁷⁰ ergy conversion factor, *k*, can be formulated as $k = (152.5 - E_t \text{MeV})/l$, where E_t is the muon $_{172}$ energy loss in the target. The μ^+ path length in ¹⁷³ the target was obtained by connecting the CsI(Tl) ¹⁷⁴ center of the μ^+ hit module and the K^+ vertex po-¹⁷⁵ sition determined by the target system. The typical k value was obtained to be 2.1–2.5 MeV⁻¹. Then, μ^+ energy spectrum from the $K_{\mu 2}$ decay is ob-¹⁷⁸ tained by taking into account the energy loss in the 179 target as $E = kl + E_t$, as shown in Fig. 5. The ¹⁸⁰ red and blue spectra indicate the calibrated energy ¹⁸¹ spectrum with and without the target energy cor-¹⁸² rection, respectively. The target energy correction ¹⁸³ improved the energy resolution to $\sigma = 2.63\%$ from ¹⁸⁴ 4.73%.

 Also, the CsI(Tl) timing information was checked ¹⁸⁶ by requiring the e^+ signals to reduce the effects from accidental backgrounds. The 40 ns clock tim- ing uncertainty of VF48 was corrected for by mea- suring the trigger signal in the same VF48 mod- ule. The timing resolution was determined to be 10.7 ± 0.1 ns (σ) .

¹⁹² **5. A new method of energy calibration using** ¹⁹³ **stopped comic-ray muons**

¹⁹⁴ It is possible to consider a new CsI(Tl) calibra-¹⁹⁵ tion method using stopped cosmic-ray muons with the subsequent e^+ emission in the CsI(Tl) calorime-¹⁹⁷ ter. This method is proposed to measure the e^+

Fig. 5: The calibrated energy spectra obtained using the $K^+ \rightarrow \mu^+ \nu_\mu$ decays. The red spectrum includes a correction for the energy loss in the target. The red lines are the fitting results assuming a Gaussian function.

 energy spectrum for a rough CsI(Tl) energy cali- bration without the need for a radioactive beam. \sin Since the maximum e^+ energy from the muon de- cay is 52*.*32 MeV, the energy calibration can be $_{202}$ performed by measuring the e^+ energy after the cosmic-ray muon stops in the CsI(Tl) crystal. The cosmic muons stop homogeneously in the CsI(Tl), and we do not need to consider the specific struc-ture of the CsI(Tl) calorimeter.

²⁰⁷ The energy distribution of the decomposed sec-²⁰⁸ ond pulse is shown in Fig. 6 as indicated by the ²⁰⁹ black dots. Here the calibration parameters ob-210 tained from the $K_{\mu2}$ decays were used. The blue ²¹¹ squares and the red open circles are the calculated e^+ and e^- energy distribution, respectively, using ²¹³ a Monte Carlo simulation based on a GEANT4 ²¹⁴ code. Electromagnetic shower leakage from the ²¹⁵ muon stopped module was taken into account. The ²¹⁶ red cross-hatched area and the green filled area are e^+ and e^- energy distributions assuming the muon ²¹⁸ yield ratio $F_{+}/F_{-}=1.1$ and 1.6, respectively [11– ²¹⁹ 15. The energy resolution of 2.63% (σ) obtained 220 from the $K_{\mu 2}$ calibration result has been used.

 $_{221}$ In order to determinate the energy calibration pa- $_{249}$ rameters using stopped cosmic-ray muons, a com- mon gain parameter relative to the energy coeffi- χ_{224} cients obtained from the $K_{\mu 2}$ calibration results was 225 introduced. The reduced χ^2_ν/NDF determined by comparing the experimental data with the simula- tion was calculated as a function of the the above relative gain coefficient, as shown in Fig. 7, where NDF is the number of degrees of freedom. The

Fig. 6: Energy distributions of *e* ⁺ (*e−*) from stopped cosmic muons. An electromagnetic gamma shower was taken into account in the simulation. The black hatched area is not used in the fitting because of the online threshold setting of 20 MeV.

230 black dots and open squares correspond to the re- 0 10 20 30 40 50 60 0 231 sults obtained by assuming $F_{+}/F_{-} = 1.1$ and 1.6, ²³² respectively. It should be noted that the fitting re-²³³ gion was chosen to be 20–60 MeV because the on-²³⁴ line energy threshold was set to 20 MeV. The black ²³⁵ hatched area was not used for the fitting. Scatter- $_{236}$ ing of the χ^2_{ν} values is due to random smearing to ²³⁷ account for the CsI(Tl) energy resolution The lines ²³⁸ in the figure represent the fitting results using a ²³⁹ parabolic function. As a result, the relative coeffi-240 cients for $F_{+}/F_{-} = 1.1$ and 1.6 were determined to
241 be 0.986 + 0.033 and 1.001 + 0.032 which indicates $be 0.986 \pm 0.033$ and 1.001 ± 0.032 , which indicates ²⁴² the gain coefficients obtained from the stopped cos-243 mic muons are consistent with those from the $K_{\mu 2}$ ²⁴⁴ events at the $3-4\%$ level. Therefore, the experimen-²⁴⁵ tal data were in good agreement with the above two ²⁴⁶ simulation models, indicating a correct understand- $\frac{1}{247}$ ing of the e^+ and e^- behavior generated from the stopped muons.

The muon lifetime curve was also measured using the time interval between the 1st and 2nd pulses, as shown in Fig. 8 by the black dots. Fitting the data with an exponential function, the decay constant was determined to be $2.06 \pm 0.03 \ \mu s$ $\left(\chi^2_{\nu}/NDF = 69.5/43\right)$, as shown by the red line. The observed time constant is a little shorter than the theoretical value which indicates that most of μ [−] events are captured by CsI nuclei and do ²⁵⁸ not contribute to the above lifetime measurement.

Fig. 7: Reduced χ^2_{ν} obtained by changing the relative gain coefficient. The black dots and open squares correspond to the results obtained by assuming $F_{+}/F_{-} = 1.1$ and 1.6, respectively. The lines in the figure represent the fitting results using a parabolic function. The gain coefficients obtained from the stopped cosmic muons are consistent with those from the $K_{\mu 2}$ events within 3–4%.

6. Conclusion

 A model function for the waveform analysis of CsI(Tl) calorimeter in the E36 experiment has been developed, and the information of the decomposed second pulses can be used for the event selection. The CsI(Tl) energy calibration was successfully 265 performed by choosing the $K_{\mu 2}$ events and impos- ing the existence of the second pulses, and the S/N $_{267}$ ratio was significantly improved. Then, the CsI(Tl) performance was carefully checked by studying the energy and timing resolutions.

 A new energy calibration method using stopped cosmic muons is proposed. The energy and timing $_{272}$ of the delayed e^+ (e^-) signals were determined by the decomposed second pulse in the double-pulse waveform analysis. The observed energy spectrum is consistent with the simulation calculation with an accuracy of $3-4\%$, indicating the establishment of a new calibration method without using any ac-celerator facilities.

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Fig. 8: Time interval between cosmic muons and the delayed e^+ (e^-) signals through the $\mu \to e\nu\bar{\nu}$ decays. The black dots and the red line are the data and fitting function, respectively.

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