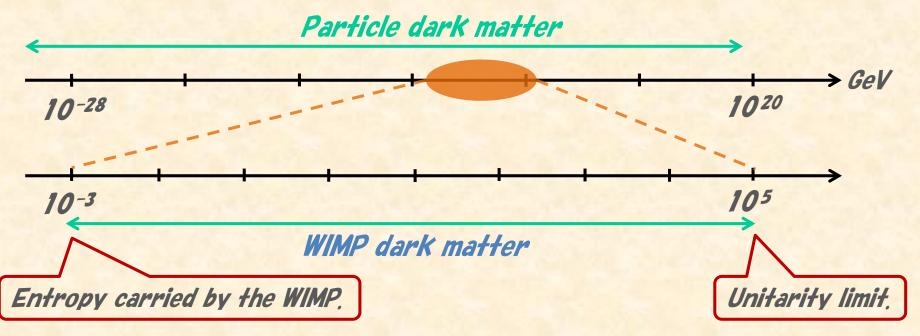
Studying WIMP without prejudice

WIMP is involved in various new physics scenarios of the EWSB.
 WIMP is predicted to be the lightest particle among new particles.
 WIMP can be searched for by not only colliders but also other ones.
 Since the properties of the WIMP is unKnown, the study of the WIMP without depending on any specific new physics models is mandatory!

• WIMP mass



Studying WIMP without prejudice

• WIMP spin

To have strong enough couplings with SM particles, WIMP is expected to have renormalizable interactions, so that its spin is 0, ½, or 1.

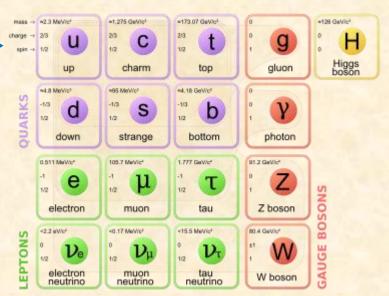
Interaction

WIMP interaction

Which interaction exists between WIMP and SM?

Classifying WIMPs by each interaction is not useful due to the consistency of FT.

Classifying WIMPs based on its quantum number is more useful for our purpose. Weak charge plays an important role!!!



WIMPs can be classified into the following three categories,

DM

WIMP

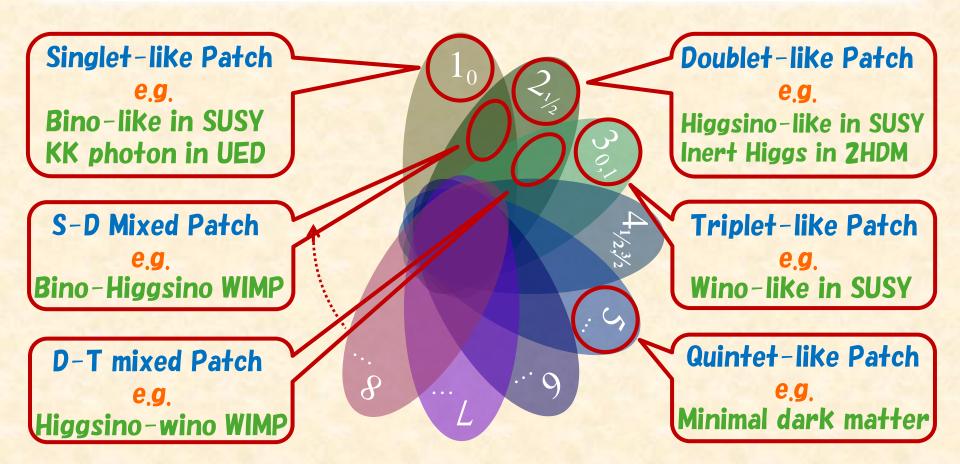
✓ WIMP has a weak charge of (almost) zero, … Singlet(-like) WIMP

- ✓ WIMP has a weak charge close of (half) integer, … EWIMP
- ✓ WIMP has a mixed weak charge due to EWSB, … Well-tempered WIMP

Studying WIMP without prejudice

After fixing its spin, the WIMP field is written by a linear combination of colorless rep, of $SU(2)_L \times U(1)_Y$ involving a EM neutral component:

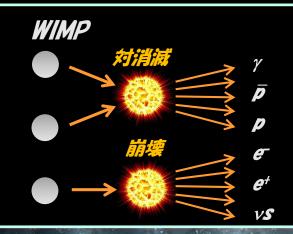
WIMP $(x) = \sum_{i} z_i [\chi_i(x)]_{\text{N.C.}}$ with $\sum_{i} |z_i|^2 = 1$





Earth

Indirect detection



Colliders

CMS

ALICE

ATLAS

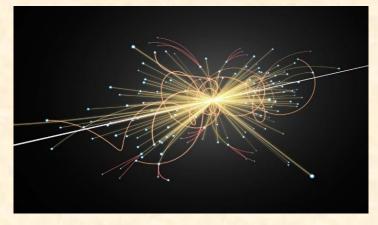
 Utilizing existent dark matter or producing it,
 The conclusive evidence of DM detection so far,

Direct detection



[CLARK planetarium website \$4]





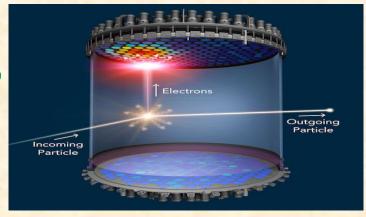
@ Colliders

WIMP is expected to be directly produced at colliders, if its energy is high enough. Hadron Collider: Interaction with quarks. Lepton Collider: Interaction with leptons.

@ Direct detection

WIMP can be detected by observing release energy by the scattering off a nucleus.

SI scattering: Int. with quarks & Higgs. SD scattering: Int. with quarks & Z boson.





@ Indirect detection

WIMP could be searched for by observing annihilation products produced at DM halo. Gamma ray: Int. with all the SM particles Cosmic ray: Int. with all the SM particles



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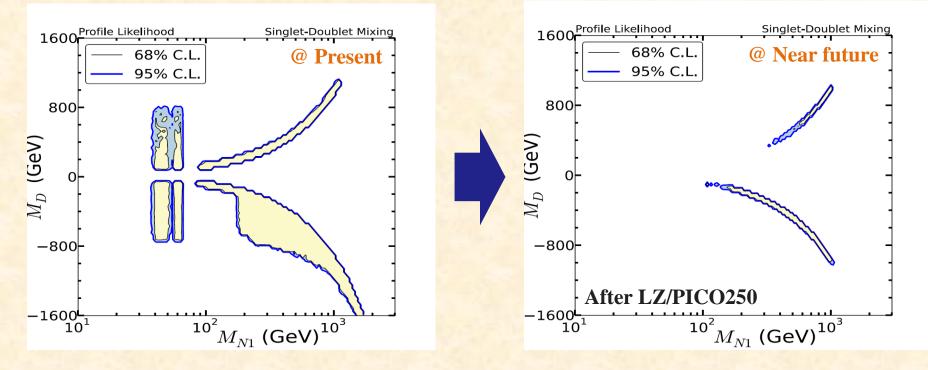
Simplest example = (Fermionic) singlet-doublet WIMP
 ✓ Typical WIMP in the traditional natural SUSY,
 ✓ Minimal contents: 1₀, 2_{1/2}, 2_{-1/2}, (Anomaly cancel,)
 ✓ 3 neutral Majorana and 1 charged Dirac fermions.

> Lagrangian assuming Z₂ symmetry making the WIMP stable is

$$\mathscr{L}_{SD} = \mathscr{L}_{kin} - \left[\frac{1}{2}M_SSS + M_DD_1 \cdot D_2 + y_1SD_1 \cdot \tilde{H} + y_2SD_2 \cdot H + \text{H.c.}\right]$$

> Scanning parameter space using MCMC to clarify the current status and future prospects of the WIMP, assuming $|y_i| \le 1$.





Direct detection is very powerful to explore the well-tempered WIMP!

The same conclusion is obtained for the most of well-tempered WIMPs, for the origin of the mixing and DM-DM-h(Z) couplings are the same. Big direct dark matter detection will be playing an important role!!!!!

EW charged WIMP (EWIMP)

Simplest example = (Fermionic) triplet-like WIMP ✓ Predicted in High-scale SUSY and MPP scenarios, ✓ Minimal contents: 3₀, (Just one representation,) ✓ 1 neutral Majorana and 1 charged Dirac fermions,

> Lagrangian assuming Z_2 symmetry making the WIMP stable is

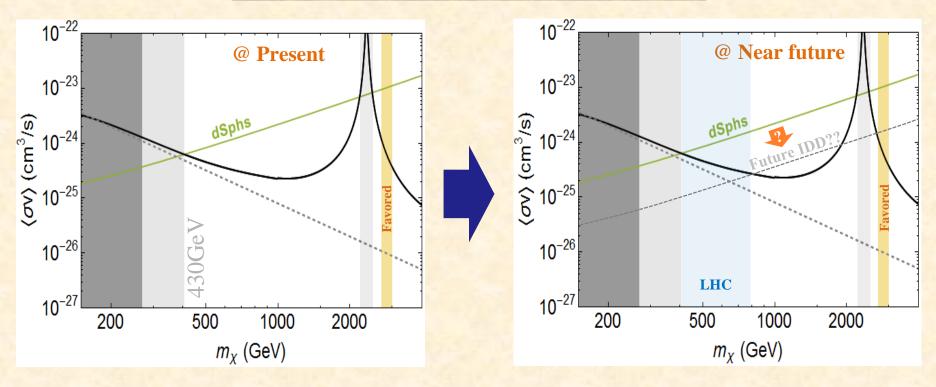
$$\mathcal{L} = \mathcal{L}_{\rm SM} + \frac{1}{2}\bar{T}\left(\not\!\!\!D - M_T\right)T$$

> Parameter space is simply defined by only one parameter M_T .

> Scanning parameter space is simple because of one parameter,

It is possible to include higher dimensional operators to take new physics effects beyond the WIMP into account, however, those do not play important roles at WIMP's phenomenology.

EW charged WIMP (EWIMP)



The WIMP seems difficult to be detected at DD searches in near future. $[\sigma_{SI} \sim 2 \ 10^{-11} \text{ pb}, \text{ and, in addition, it may be cancelled by BSM contributions.}]$ LHC will explore the WIMP mass region below 500GeV. Can it go more? IDD searches are promising, for the WIMP's annihilation is enhanced!!! [The enhancement is from the Sommerfeld effect, Hisano, S.M., Nojiri, 2014.] γ -ray obs. (Fermi, CTA) \rightarrow IDD (γ from dSphs) \leftarrow DM dist. (PSC, PFS)

Singlet-like WIMP (Heavy mediator)

Simplest example = (Fermionic) singlet - like WIMP
✓ Predicted in all of the DS scenarios involving WIMP.
✓ Minimal contents: 1₀ + Mediator ★\$!
✓ 1 neutral Majorana and mediator states(s).

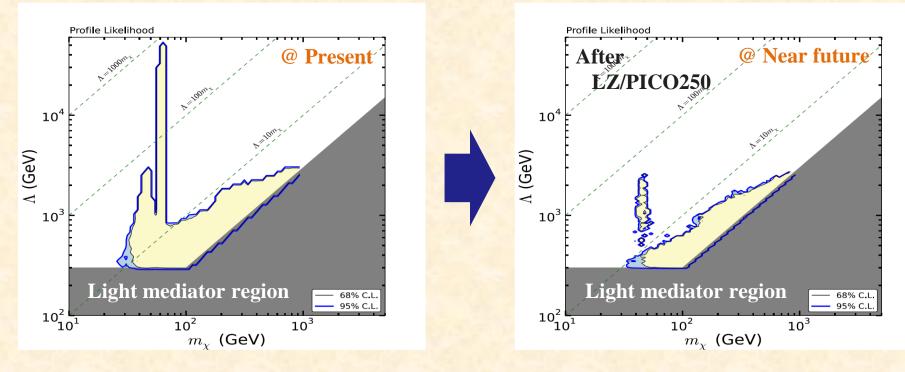
When the mediator is heavier enough than the WIMP and the EW scale, the phenomenology is effectively described by the EFT,

 $\mathcal{L}_{\rm EFT} \supset \frac{c_S}{2\Lambda} (\bar{\chi}\chi) |H|^2 + \frac{c_P}{2\Lambda} (\bar{\chi}i\gamma_5\chi) |H|^2 + \sum_f \frac{c_f}{2\Lambda^2} (\bar{\chi}\gamma^\mu\gamma_5\chi) (\bar{f}\gamma_\mu f) + \frac{c_H}{2\Lambda^2} (\bar{\chi}\gamma^\mu\gamma_5\chi) (H^\dagger i\overleftrightarrow{D_\mu} H)$

where A represents the typical mass scale of the mediator.

- > Parameter space is very complicated, ³ about 10 parameters.
- > Scanning parameter space using MCMC, assuming CP invariance and the flavor blindness of the WIMP interaction with $|c_i| \leq 1$.

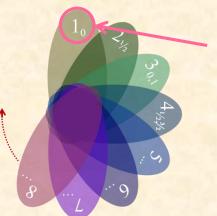
Singlet-like WIMP (Heavy mediator)



Direct detection is powerful to explore the H- & Z-resonance regions. The four Fermi interactions governs the other region with $\Lambda < 10m_{DM}$, [This region is not so much searched for at DD and LHC exps in near future!]

[It is governed mainly by the interactions with leptons or Z-boson.]

Light singlet WIMP (Light mediator)



Simplest example = (Fermionic) singlet-like WIMP ✓ Predicted in all of the DS scenarios involving WIMP. ✓ Minimal contents: 1₀ + Scalar/Vector Mediator, ✓ 1 neutral Majorana and mediator states,

Let us consider the case of a light singlet WIMP + a scalar mediator!

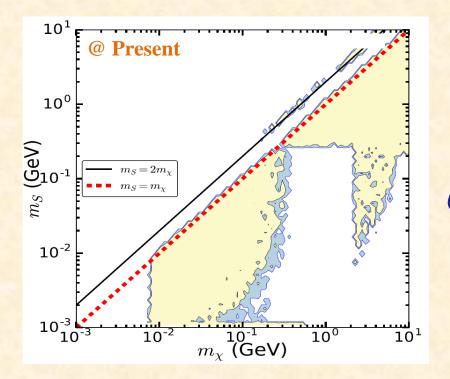
> Lagrangian involving all possible renormalizable interactions:

$$\mathcal{L} = \mathcal{L}_{\rm SM} + \frac{1}{2}\bar{\chi}(i\partial \!\!\!/ - m_{\chi})\chi + \frac{1}{2}(\partial \phi)^2 - \frac{c_s}{2}\phi\bar{\chi}\chi - \frac{c_p}{2}i\phi\bar{\chi}\gamma^5\chi - V(\phi, H),$$

> Parameter space is again very complicated, ³8 parameters,

> Scanning parameter space using MCMC, assuming CP invariance $(c_p = 0)$ with being dimension-less(full) coupling $\leq 1(1TeV)$.

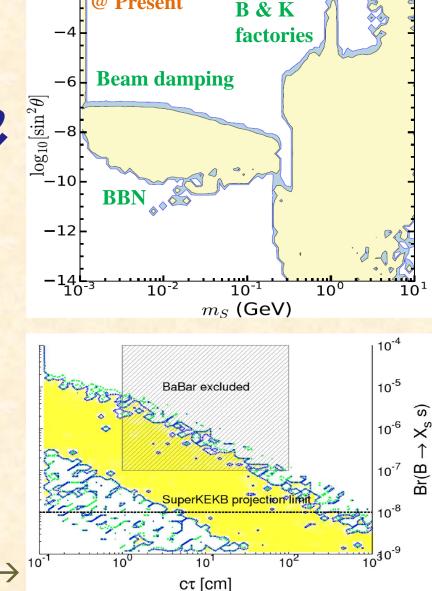
Light singlet WIMP (Light mediator)



Various experiments will contribute in order to explore the light WIMP.

Search for the light mediator s will play a crucial role in this program. [An exception is direct DM detection, where the WIMP plays a central role.]

Long-lived particle search @ Belle ||



@ Present



- ✓ We have discussed WIMP candidates and new physics of EWSB!
- ✓ Lot of new physics models predict the existence of WIMP:
 - Traditional natural SUSY... Well-tempered WIMPFocus point SUSY scenario... Doublet-like WIMPHigh-scale SUSY scenario... Triplet-like WIMPMultiple point principle... Triplet-like WIMPDark sector scenario... Singlet-like WIMP
- Current status & future prospect to search for the WIMP are Direct detection searches are (and will be) playing a very important role to search for the well-tempered WIMP. Indirect detection searches will be the only way to explore the electroweakly charged WIMP (EWIMP) in near future. Leptophilic and Z-pole regions will remain unexplored for the singlet-like WIMP with heavy Mediator particle(s). Many studies are now on-going for the singlet-like WIMP with light Mediator(s). Among those, the light WIMP region is interest, which will be explored by e.g. Belle II.